

1 Evolving magma temperature and volatile contents
2 over the 2008-2018 summit eruption of Kīlauea
3 Volcano

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5 **Magma rheology and volatile contents exert primary and highly nonlinear**
6 **controls on volcanic activity. Subtle changes in these magma properties can**
7 **modulate eruption style and hazards, making in situ inference of their tem-**
8 **poral evolution vital for volcano monitoring. Here we study thousands of im-**
9 **pulsive magma oscillations within the shallow conduit and lava lake of Kīlauea**
10 **Volcano, Hawai‘i, USA over the 2008-2018 summit eruptive sequence, encoded**
11 **by ‘Very-Long-Period’ seismic events and ground deformation. Inversion of**
12 **these data with a petrologically informed model of magma dynamics reveals**
13 **significant variation in temperature and highly disequilibrium volatile con-**
14 **tents over days to years, within a transport network that evolved over the**
15 **eruption. Our results suggest a framework for inferring subsurface magma**
16 **dynamics associated with prolonged eruptions in near real time that synthe-**
17 **sizes petrologic and geophysical volcano monitoring approaches.**

18 **One-Sentence Summary**

19 Resonant magma oscillations reveal evolving magma properties over a decade long eruption
20 at Kīlauea Volcano, Hawai‘i, USA.

23 Introduction

24 Kīlauea volcano, Hawai‘i, USA, is one of the most active, best monitored, and best stud-
25 ied volcanoes on Earth (1), serving as a focal point for volcanologic research (2). However,
26 resolving in situ variation in subsurface magma dynamics remains a challenge at Kīlauea and
27 volcanoes globally (3). The 2008-2018 Kīlauea summit eruption represents an opportunity to
28 address this knowledge gap. The eruption involved a persistent lava lake in the Halema‘uma‘u
29 summit vent and multiple subsurface magma intrusions and East Rift Zone eruptions, ending
30 with a spectacular caldera collapse sequence representing the highest historical sustained erup-
31 tion rate at Kīlauea (4–6). Previous studies suggested the main Kīlauea shallow summit magma
32 plumbing system during this time consisted of the 1-2 km deep Halema‘uma‘u reservoir and
33 the 3-5 km deep South Caldera reservoir (Fig. 1) (7, 8). The Halema‘uma‘u reservoir and over-
34 lying lava lake were continuously connected (4) by a ~ 10 m wide conduit (9). Magma passed
35 through the summit en route to the East Rift Zone, although the nature of hydraulic connections
36 between the summit reservoirs, rift zone, and deeper magma sources are not well known (8, 10).

37 A wide range of data, interpreted using physical and chemical models, inform this picture
38 of magma dynamics. Transport geometry is constrained primarily through inversion of seismic
39 and geodetic data (7, 9, 11). Continuous gravity data are only available over limited time seg-
40 ments, but constrain the density of magma in the lava lake and suggest temporal variation of up
41 to 1500 kg/m^3 (12). Analysis of erupted products provides limited temporal and spatial resolu-
42 tion, but suggests Halema‘uma‘u magma consists of near-liquidus ($1150\text{-}1300 \text{ }^\circ\text{C}$) crystal-poor
43 basalt outgassed in CO_2 with respect to the primary mantle magma (13, 14). Subsurface magma
44 volatile contents are also indirectly informed by continuous gas emissions (13, 15, 16). These
45 analyses suggest significant disequilibrium outgassing, or mechanical decoupling of gas bub-
46 bles from melt due to continuous convecting and outgassing (17). However, geochemical and
47 geophysical data are rarely combined in a quantitative manner.

48 Very-Long-Period (VLP) seismicity, with energy concentrated at periods above 2 s, has the
49 potential to help unify these diverse constraints. VLP seismicity is prevalent at many volcanoes
50 and often inferred to represent transient magma flow (18), thus directly probing magma proper-
51 ties and transport geometry in ways not readily obtainable by other geophysical analyses. VLP
52 signals are part of a spectrum of oscillatory motions that can result from impulsive or contin-
53 uous forcing of magma transport structures (19, 20), but the VLP band is advantageous due to
54 being less sensitive to path distortions from heterogeneous earth structure than shorter period
55 signals.

56 Multiple resonant modes have been identified at Kīlauea, but the dominant VLP signal is
57 from ‘conduit-reservoir’ resonance in which stratified magma in the conduit and lake sloshes
58 in and out of the underlying reservoir (9, 21, 22) (Fig. 1). This resonance occurs sometimes
59 as continuous tremor but most often as discrete minutes-long events triggered both from the
60 lake surface (such as via rockfalls from the crater walls) and from depth (22, 23). Oscillation
61 restoring forces are from gravity and magma reservoir elasticity, while damping is from viscous

62 drag on the conduit walls. Resonant period is primarily sensitive to conduit length and bulk
63 magma density/density stratification (9). Decay rate is quantified by quality factor (the ratio
64 of energy stored per cycle over energy lost per cycle) and is primarily sensitive to conduit ra-
65 dius and apparent magma viscosity. In the shallow Halema‘uma‘u magma system, where melt
66 composition doesn’t vary much in time or space and where crystal contents are low (13, 14, 24),
67 magma density is primarily controlled by porosity and magma viscosity is primarily controlled
68 by porosity and temperature. In chemical equilibrium, gas mass fraction (hence porosity) de-
69 pends upon total volatile mass fraction and pressure-dependent solubility of dominant volatile
70 species (H_2O , CO_2 , and sulfur) (25).

71 VLP seismicity at Kīlauea thus reflects evolving magma thermal and chemical state as well
72 as transport structures. Over the 2008-2018 Kīlauea eruption, thousands of conduit-reservoir
73 resonance events provide an unprecedented record of time-evolving subsurface magma trans-
74 port.

75 **Approach: Inferring magma properties from geophysical data**

76 Fig. 2 outlines our workflow. We first conduct kinematic elastic inversions between 2008-
77 2018 of continuous Global Navigation Satellite System (GNSS) ground deformation data (26)
78 (Fig. S3, S4) for shallow magma reservoir pressure histories. In particular, Halema‘uma‘u
79 reservoir pressure constrains magma column density in the overlying summit lava lake and
80 conduit. Summit deformation at Kīlauea is complex: to resolve Halema‘uma‘u reservoir pres-
81 sure we build on constraints from previous geodetic studies (7, 11, 27) and include three known
82 deformation sources (26).

83 We next use a perturbation approach to model transient flow associated with conduit-reservoir
84 magma resonance (26) (Fig. 1), extending previous analyses (9, 21). We treat fluid properties
85 of the multi-phase magma as functions of magmastic pressure (an approximation given slow
86 exchange flow within the conduit/lava lake (28)), temperature, and vertically stratified total
87 volatile mass fractions (CO_2+H_2O , Fig. 1, S2), neglecting crystals and assuming an average
88 melt composition based on 2008-2010 Halema‘uma‘u samples (13, 25, 29–31). We use this
89 model to invert for magma properties from Halema‘uma‘u reservoir pressure, lava lake eleva-
90 tion and areal extent (4, 32), and the resonant period and quality factor of VLP seismic events
91 cataloged over 2008-2018 by (22) (Fig. 2) (26).

92 Resolving time evolution of shallow magma properties at Kīlauea is a long-standing chal-
93 lenge (9, 33, 34). We focus on shorter term changes in multiphase magma properties by assum-
94 ing a fixed magma system geometry based on previous inversions (7, 9, 11). Four additional
95 assumptions are made to facilitate unique inversions for magma properties (26) (supplementary
96 text): 1. Temperature is spatially uniform in the conduit and lake. This is justified because
97 the conduit undergoes quasi-steady exchange flow/mixing (35) and the lake contributes negli-
98 gibly to viscous damping. 2. Magma in the conduit/lake has a fixed total (dissolved+exsolved)
99 H_2O/CO_2 mass ratio. Volatile composition could vary over time, but is unconstrained in our
100 model without additional data so we fix volatile ratios based on erupted products and gas emis-

101 sions (13, 14, 36). 3. Total volatile mass fraction varies linearly with depth (Fig. 1) subject to
102 stable stratification which should be approximately valid for the largely quiescent magma col-
103 umn. 4. Total volatile mass fraction at the lake surface is constant. While there is known to be
104 some variation in porosity near the lake surface from continuous gravity data (37), these data are
105 not available over most of the timespan. Additionally, our model exhibits minimal sensitivity to
106 density stratification within the lake; it is primarily sensitive to average density (which controls
107 the magmatic pressure load of the lake on the conduit).

108 We test different fixed parameter combinations and conduct an a posteriori assessment of
109 these assumptions. The magma properties we invert for are: 1. magma temperature, 2. conduit
110 average total volatile mass fraction X^{avg} , and 3. total volatile mass fraction stratification (differ-
111 ence between conduit top and conduit bottom) ΔX . We note that while the magma temperature
112 parameter is applied to the whole magma column, the model is primarily sensitive to conduit
113 temperature. We also note that due to tradeoffs between volatile contents at the bottom and top
114 of the lava lake, ΔX should be considered to represent a general volatile stratification over the
115 whole magma column (conduit and lava lake).

116 Results

117 For our reference fixed parameters, Fig. 3 shows the timeline of GNSS inversion results
118 and VLP magma resonance inversion results, along with other data. Shaded regions in Fig.
119 3 show the envelope of inversion results obtained by varying individual fixed parameters over
120 feasible ranges, as detailed in the supplementary text (Fig. S5). Evolution of magma system
121 geometry, which is not considered in our inversions, is more likely to affect trends in inversion
122 results over long (year or more) timescales. In particular, inversion results with the reference
123 fixed parameters are likely not reliable in 2009-early 2010 and mid 2011 (discussion). On short
124 timescales, noise in input data likely contributes to scatter and outliers in the inversion results.
125 We thus focus most analysis on temporally averaged values, and in particular on the relative
126 variability in these values over timescales of a year or less rather than their absolute value at a
127 given time. Fig. 4 shows amplitude spectra, coherence, and phase lags between data sets with
128 95% significance thresholds (supplementary text). Additional analyses are shown in Fig. S6,
129 S7, and S8.

130 As expected for an open-vent magma system, Halema‘uma‘u reservoir pressure is well
131 correlated with lava lake elevation over timescales from days to about a year (Fig. 3, 4)
132 (4, 22). Strong coherence between Halema‘uma‘u and South Caldera reservoir pressures over
133 timescales of days to months (Fig. 4, S6) suggests that magma is often transferred between the
134 reservoirs, although the anticorrelation implies hydraulic disequilibrium. This could indicate an
135 intermittent connection, consistent with the unsteady connectivity inferred during hours-days
136 long “deflation-inflation” events (6, 8, 38). We are not aware of any other settings where a
137 consistent anticorrelation is observed between different magma reservoirs at the same volcano,
138 although intermittent hydraulic connections have been inferred between Kīlauea and Mauna
139 Loa (39), as well as at other volcanoes such as Soufrière Hills (40) and Etna (41).

140 Different fixed parameters affect the absolute value of inverted magma temperature, but the
141 pattern of relative temporal variation is robust and the magnitude of such changes varies by less
142 than ~ 20 °C (Fig. 3, S5). Inverted temperature is primarily sensitive to conduit radius; decreas-
143 ing radius by 10 m (to 5 m) uniformly increases temperatures by ~ 60 °C while increasing radius
144 by 10 m (to 25 m) uniformly decreases temperatures by ~ 40 °C. Conduit magma temperatures
145 span the full 1150-1300 °C range of Halema‘uma‘u magma storage temperatures previously es-
146 timated from ejecta geothermometry (13, 24), although it is difficult to make a direct comparison
147 given uncertainty in the depths and/or timescales recorded by geothermometers.

148 On timescales from days-months, temperature exhibits up to 100 °C variation (Fig. 3),
149 corresponding to up to an order of magnitude variation in magma viscosity (Fig. S2, S8).
150 Temperature and resonant quality factor are strongly correlated (Fig. S6), which suggests that
151 temperature is a primary driver of variations in magma viscosity. The dominance of temperature
152 is unexpected because porosity has previously been proposed as a likely source of variation
153 in VLP quality factor (21) and is known to vary significantly as bubbles rise and accumulate
154 (28, 37).

155 Different fixed parameters affect the inverted absolute value of X^{avg} by up to ~ 1 wt%, but
156 the pattern of relative temporal variation is robust and the magnitude of such changes varies
157 by less than ~ 0.4 wt% (Fig. 3, S5). Similarly, different fixed parameters affect the inverted
158 absolute value of ΔX by up to ~ 1 wt%, but the pattern of relative temporal variation is robust
159 and the magnitude of such changes varies by less than ~ 0.2 wt% (Fig. 3, S5). Over most of the
160 timeline X^{avg} is greater than the inferred primary magma volatile mass fraction of 1-2 wt%, a
161 notable accumulation particularly since some of the primary CO₂ may have already been lost
162 at depth (14, 36, 42). Additionally, ΔX is mostly similar to or larger than inferred primary
163 magma volatile mass fraction. Together these indicate significant departures from equilibrium
164 outgassing, with an accumulation of volatiles in the upper conduit and lava lake.

165 On timescales of days-months, X^{avg} varies by up to ~ 0.6 wt% and ΔX varies by up to
166 ~ 1 wt% (Fig. 3). That this temporal variation is similar to the inferred primary magma’s
167 total volatile mass fraction of 1-2 wt% (36, 42) suggests significant variations in the outgassing
168 regime (14). The only volatile species with continuous emission measurements that can be
169 compared with ΔX and X^{avg} is SO₂. SO₂ has roughly similar solubility to H₂O in mafic
170 melts (43) and so will approximately trade-off with H₂O in our model. SO₂ emissions exhibit
171 strong variation (an order of magnitude or more) on timescales from days to years (15, 16). We
172 do not observe consistently strong coherence between ΔX or X^{avg} and SO₂ emissions (Fig.
173 S6), although several pronounced increases in either ΔX or X^{avg} do correspond to increases in
174 SO₂ (e.g., Apr 2015, Jan 2016, Oct 2016, and Aug 2017). Inconsistent coherence could partly
175 reflect the high uncertainty in SO₂ emission data, although we note that gas emissions from the
176 lava lake surface will not necessarily directly correlate with the amount of volatiles accumulated
177 in the magma column. In fact, the strong in-phase coherence between Halema‘uma‘u reservoir
178 pressure (or lava lake elevation) and ΔX on timescales of less than 90 days (Fig. 4) suggests
179 that volatiles build up in the upper conduit/lake as magma accumulates in the Halema‘uma‘u

180 system, rather than maintaining a steady volatile mass balance through the shallow magma
181 column. This could reflect an increase in volatile flux (e.g., from magma recharge), but could
182 also be caused by less efficient outgassing through the lava lake as it fills.

183 **Discussion**

184 Halema‘uma‘u magma mass balance

185 Maintaining a persistent lava lake for a decade requires a remarkable thermal and mechan-
186 ical balance. Relatively constant magma supply from depth is needed to drive continuous con-
187 vection, but supply must be countered by sufficient outflux to prevent conditions leading to
188 violent eruption. Ground deformation and VLP seismicity provide a quasi-continuous probe
189 of magma properties that facilitates interrogation of the multiscale processes maintaining (and
190 modulating) this balance within the Halema‘uma‘u reservoir during an extended eruption.

191 In general, magma reservoir pressure can change even without any magma input due to
192 gas exsolution and (to a lesser extent) crystallization. However, since the low-viscosity mafic
193 melt and open-vent structure of Halema‘uma‘u facilitates gas escape, reservoir pressurization
194 has been inferred to reflect accumulation of melt either due to changes in influx (e.g., recharge
195 from the South Caldera reservoir or deeper storage regions) or outflux (e.g., to the East Rift
196 Zone) (4, 44). For example, the inferred causes of the May 2015 summit intrusion, the 2018
197 eruption, and the prevalent hours-days long “deflation-inflation” summit deformation events
198 are: months of increased magma influx (4, 6, 27), months of reduced magma outflux (45), and
199 transient restrictions of magma influx or outflux (6, 8, 38). However, the general controls on
200 magma mass balance over days-years are unknown. The 60 and 130 day period spectral peaks
201 in Halema‘uma‘u reservoir pressure (also apparent in temperature, ΔX , and X^{avg}) (Fig. 4) may
202 indicate dominant timescales for such changes in influx-outflux (4). Quasi-periodic deformation
203 and/or eruptive activity on similar timescales has also been observed at other volcanoes (46, 47).

204 We might expect magma recharge to increase conduit temperature, although this would de-
205 pend on the temperature and influx of recharging magma, and also its path through the $\sim 4 \text{ km}^3$
206 of near-liquidus magma in the Halema‘uma‘u reservoir (11, 24). The inferred 2011-2012 aver-
207 age magma supply rate of $\sim 10^9 \text{ kg/day}$ (34) would permit complete exchange with the $\sim 10^{10} \text{ kg}$
208 of magma in the conduit and lava lake over a week. However, if this injected magma were uni-
209 formly mixed with the magma in the reservoir ($\sim 10^{13} \text{ kg}$ assuming a density of 2500 kg/m^3) at
210 a 100°C temperature difference, the mixture temperature would only increase by $\sim 0.01^\circ\text{C/day}$
211 (neglecting latent heat and outflow). Given the poor coherence between Halema‘uma‘u reser-
212 voir pressure (or lava lake elevation) and temperature (Fig. 4), we expect that melt injected into
213 the Halema‘uma‘u reservoir generally either was not appreciably hotter than existing magma
214 and/or was not directly routed to the conduit.

215 One prominent exception that could exemplify an influx of hotter melt from depth is the
216 persistent $\sim 100^\circ\text{C}$ increase in temperature six months before the Mar 2011 Kamoamoia fissure
217 eruption. There was no corresponding increase in volatile mass fractions, potentially due to

218 deeper separation and upward flux of volatiles over the preceding months of elevated volatile
219 mass fractions. Interestingly, temperature then dropped by ~ 100 °C in the months leading
220 up to the eruption, which we expect relates to lava lake downwelling rather than magma in-
221 flux/outflux, as discussed in the next section. Another potential example of hot melt influx is
222 the ~ 90 °C increase in temperature between the May 2012 slow-slip event on Kīlauea’s south
223 flank décollement and the Oct 2012 intrusion, although there was also no corresponding in-
224 crease in volatile mass fractions. The temperature increase supports previous suggestions that
225 slow-slip events are linked to magmatism (48), although we do not see similar temperature
226 increases immediately following the 2010 or 2015 slow-slip events.

227 It is less obvious what changes in magma properties might be expected from decreased
228 magma outflux, so we use the 2018 eruption as a case study. The months of pressurization
229 preceding the eruption are accompanied by a decrease in magma temperature and increase in
230 X^{avg} , but these do not clearly stand out from the background variation over the preceding year
231 (Fig. 3). The lack of clear changes in magma properties is consistent with the idea that the 2018
232 eruption was triggered by decreasing outflux rather than by recharge (45) and, by extension,
233 suggests that outflux does not necessarily drive significant changes in shallow magma proper-
234 ties. The May 2014 and May 2015 intrusions were also preceded by a month of Halema‘uma‘u
235 reservoir pressurization without other clearly associated changes in the summit magma system.
236 The lack of clear changes in magma properties would seem to suggest they were induced by de-
237 creased magma outflux, although at least in 2015 changes in East Rift Zone lava effusion were
238 not apparent (4, 6, 27). The Jun 2014 and May 2016 Pu‘u‘Ō‘ō vent openings were not preceded
239 by significant pressurization of the shallow summit magma system, suggesting they were not
240 primarily caused by increased melt flux from the summit but rather by processes along the rift
241 zone.

242 Shallow magma dynamics

243 Our results illuminate shallow fluid dynamic processes underlying a persistent lava lake.
244 Observed covariation of parameters in our inversions suggest that volatile mass fraction and
245 temperature in the conduit and lava lake vary in ways not always directly related to Halema‘uma‘u
246 reservoir magma influx/outflux. We infer that such variation occurs due to unsteady exchange
247 flow between the conduit and Halema‘uma‘u reservoir (49), as well as due to changing convec-
248 tive efficiency in the lava lake and/or surface crust dynamics (which influence the outgassing
249 rate and efficiency of heat loss to the atmosphere and host rock) (4, 50).

250 The negative correlation on timescales of months or less between X^{avg} and temperature
251 (Fig. 4, S6) likely reflects such dynamics, since relatively poor coherence with Halema‘uma‘u
252 reservoir pressure (or lava lake elevation) indicates neither X^{avg} or temperature are primarily
253 driven by magma mass balance. Simple thermal arguments suggest likely causes of tempera-
254 ture variation. Atmospheric heat exchange at the lake surface will be dominated by radiative
255 heat flux $\phi_r = A\epsilon\sigma(T_{surf}^4 - T_{atm}^4)$, where ϕ_r is ~ 1 GW for lake surface area $A \approx 10^4$ m²,
256 thermal emissivity $\epsilon \approx 0.8$, Stefan-Boltzmann constant $\sigma = 5.7 \times 10^{-8}$ Wm⁻²K⁻⁴, and average

257 surface temperature $T_s \approx 700$ °C (50). Heat flux to the host rock depends upon hydrothermal
 258 circulation, but can be approximated with an effective thermal conductivity $\phi_c = k_e \Delta T / \Delta L$,
 259 where ϕ_c is 10-1000 W/m² for k_e of 2-20 Wm⁻¹C⁻¹ (51) and temperature gradient $\Delta T / \Delta L$ of
 260 10-100 °C/m (52). Total heat transfer rate Φ from the conduit and lake (surface area $\sim 10^5$ m²)
 261 and from the Halema‘uma‘u reservoir (surface area $\sim 10^7$ m²) is 1-100 MW and 0.1-10 GW,
 262 respectively. Neglecting latent heat, average temperature of a magma mass M will decrease
 263 as $dT/dt = \Phi / (c_p M)$. For specific heat $c_p \approx 1000$ Jkg⁻¹K⁻¹, average temperature of the
 264 $\sim 10^{10}$ kg of magma in the conduit and lake could decrease by ~ 10 °C/day, whereas average
 265 temperature of the $\sim 10^{13}$ kg of magma in the Halema‘uma‘u reservoir would only decrease
 266 by ~ 0.01 -1 °C/month. We thus expect the prevalent temperature drops of 100 °C or more that
 267 occur over days-weeks represent downwelling of magma that cooled in the upper lava lake.
 268 Episodic downwelling suggests episodically decoupled convection cells in the lava lake, rather
 269 than a convective regime that settles persistently into one of the configurations previously pro-
 270 posed (6, 53). This mechanism likely explains the ~ 100 °C temperature drop preceding the Mar
 271 2011 Kamoamoia fissure eruption, where a changing convective regime is perhaps related to the
 272 rapidly filling lava lake and/or high short-term (hours-days) variability in lava lake elevation
 273 during this time. In some other cases rapid lava lake draining might also induce downwelling
 274 of cool magma. This downwelling could explain the days-long temperature decreases accom-
 275 panying the Oct 2012 and May 2014 intrusions, although if so it is interesting that the 2015
 276 intrusion did not cause a temperature drop.

277 An evolving magma plumbing system geometry

278 Given a consistent open hydraulic connection between the Halema‘uma‘u reservoir and lava
 279 lake, the weakening coherence between them over years or longer (Fig. 4) could represent ei-
 280 ther changes in the magma column density or in the relation between reservoir pressure and
 281 ground deformation (a function of geometry and poro-visco-elastic rock properties). Our fixed
 282 geometry inversions test the former and show that for a range of feasible fixed parameter values
 283 (Fig. S5) very high values of X^{avg} and/or ΔX are required over some portion of the timeline
 284 (e.g., 2009 through mid-2010 for reference parameters). These volatile contents would corre-
 285 spond to a foam in the upper conduit and lava lake with an average porosity in excess of 90%.
 286 Available constraints from gravity data (12) suggest average porosity in the lava lake of only
 287 up to 70%, so the higher values inferred at early times are likely unrealistic. We thus expect
 288 subsurface magma plumbing system geometry evolved over time, which could also contribute
 289 to the weak coherence between inverted South Caldera and Halema‘uma‘u reservoir pressures
 290 over long timescales (Fig. 4).

291 Changes in conduit length (reservoir-roof depth) of ~ 10 m or changes in conduit radius of
 292 ~ 1 m could measurably impact VLP resonance period and quality factor at Kīlauea (Fig. S5).
 293 Such changes might occur gradually due to processes such as viscous deformation of the host
 294 rock, thermal/mechanical erosion, or crystallization. Geometry could also change abruptly due
 295 to host rock failure or opening/closing of hydraulically connected dikes/sills. To fit the low
 296 VLP periods in 2009-2010 with realistic volatile contents, a ~ 100 m higher reservoir roof ele-

297 vation (510 instead of 410 m ASL, which is within estimated uncertainty (11)) and/or strongly
298 tapered conduit (e.g., top radius < 5 m and bottom radius > 15 m) is required (Fig. S9). It is
299 unlikely that the roof of an ellipsoidal reservoir would have grown downward this much over
300 year timescales due to crystallization, so it may have been shallower throughout the eruption.
301 In this case the drastic change in VLP periods over the early part of the eruption likely repre-
302 sent an evolving conduit geometry due to some combination of a widening upper conduit and a
303 change in conduit length due to a changing dip angle and reservoir attachment depth. A shallow
304 dike/sill above the main Halema'uma'u reservoir could have also impacted the resonance (54);
305 this would be potentially consistent with some seismic inversions (21, 33) but such additional
306 source complexity is not needed according to other seismic and geodetic inversions (7, 9, 11).

307 Towards a new generation of volcano monitoring

308 Resolving the dynamics of subsurface magma transport is a grand challenge that dictates
309 hazard forecasting efficacy as well as connections between active volcanic processes and the
310 geologic record. Inferring relative changes in magma properties over days to months by iden-
311 tifying the fluid origin of Very-Long-Period seismic events represents a concrete step towards
312 unifying the inversion of geophysical and geochemical data. In particular, we have resolved
313 temperature changes of over 100 °C that likely reflect both convective overturns and magma
314 recharge. We have also resolved stratified volatile profiles that represent a highly disequilibrium
315 outgassing regime. Volatile contents vary by over 1 wt% on timescales from days to months,
316 revealing an unsteady shallow volatile mass balance. We have also inferred an evolving magma
317 system geometry, highlighting the need to develop models and data sets that can deconvolve
318 changing fluid properties from changing transport pathways.

319 Incorporating additional data would yield even more precise constraints on multiphase magma
320 properties and their depth variation. For example, continuous gravity data would provide
321 independent constraints on magma density in the lake. Video of lake surface oscillations
322 could independently constrain vertical motions of the lake and triggering mechanisms of VLP
323 events. Additionally, surface gas emission data could constrain volatile stratification and out-
324 gassing/convective regimes if combined with models for gas flux through the magma column.

325 Similar Very-Long-Period events have been detected at Vanuatu and Erebus volcanoes (55,
326 56) and are expected at open-vent volcanoes generally (20), suggesting this type of analysis
327 could be adapted to improve near real time monitoring at other eruptions. These data will
328 inform basic volcano science and lead to better understanding of physical controls on volcanic
329 eruptions.

330 **Materials and Methods**

331 GNSS inversions

332 To obtain time series of pressure change in the Halema'uma'u reservoir, we must consider
333 other known sources of ground deformation at Kīlauea summit: the South Caldera reservoir
334 (7, 8), 2015 intrusion (27), and steady slip along the south flank décollement (57) (Fig. S3).

335 We assume a temporally fixed geometry for the three magma reservoirs (Fig. 1, supplementary
 336 text), but constrain the 2015 intrusion to be an active deformation source only over May 13-
 337 17 (27). We adopt the 2 km deep 4 km³ ellipsoidal Halema‘uma‘u reservoir geometry and
 338 3 GPa rock shear modulus from (11), consistent with other studies (7, 9, 10, 58, 59). We assume
 339 a horizontal centroid location of the South Caldera reservoir based on inversions of (60); depth
 340 and geometry are less well constrained so we choose a reference 20 km³ sphere centered 4 km
 341 deep and test different values based on published ranges (7, 10, 58). We fix the 2015 intrusion
 342 geometry following (27).

343 Reservoir pressures are found using linear least square inversions (supplementary text) of
 344 daily average surface position solutions from Nevada Geodetic Laboratory (61) for GNSS sta-
 345 tions within a few km of the reservoirs (Fig. 1), corrected for steady background south flank
 346 slip with the multi-component dislocation model of (57) (Fig. S3, S4). We use an approximate
 347 solution for deformation associated with a pressurized ellipsoid in an elastic half space (62) for
 348 each of the three magma bodies.

349 Conduit-reservoir magma oscillation model

350 We model VLP seismic events as small amplitude, isothermal and incompressible oscilla-
 351 tory magma flow within a lava lake-conduit-reservoir system. The model is extended from (20)
 352 to include inertial effects in the lava lake and experimentally constrained models for multiphase
 353 magma properties (supplementary text). We consider an inclined radially symmetric magma
 354 column encompassing the lava lake and conduit, underlain by a reservoir within elastic rocks
 355 (Fig. 1).

356 The magma column prior to VLP events is assumed magmastatic, justified because fluid
 357 particle velocities associated with resonance are larger than background exchange flow (20).
 358 During VLPs, viscous drag is determined from shear stress at the magma column wall where a
 359 no-slip velocity condition is enforced. With z and r distance parallel and perpendicular to the
 360 magma column axis (a function of conduit dip from horizontal θ), linearized conservation of
 361 momentum (primed variables) around a background state (bars) is

$$\frac{d\langle v' \rangle}{dt} \bar{\rho} = \langle u' \rangle \sin(\theta) \frac{d\bar{\rho}}{dz} g - \sin(\theta) \left[\frac{\partial p'}{\partial z} + \frac{2\mu}{R} \frac{\partial v'}{\partial r} \right]_R. \quad (1)$$

362 Here $\langle u' \rangle$ is cross-sectionally averaged conduit-parallel fluid particle displacement (so the ori-
 363 entation of $\langle u' \rangle$ is a function of θ), v' is conduit-parallel fluid particle velocity, $\langle v' \rangle$ is cross-
 364 sectionally averaged v' (the time derivative of $\langle u' \rangle$), ρ is magma density, p' is pressure per-
 365 turbation, μ is dynamic viscosity, and R is conduit radius. Conservation of mass is $\langle u' \rangle =$
 366 $\langle u'_0 \rangle R_0^2 / R^2$, where subscript 0 indicates evaluation at the bottom of the magma column (Fig. 1).

367 We assume equilibrium joint solubility of CO₂ and H₂O in Halema‘uma‘u composition
 368 melts (13) as a function of pressure and gas composition (25) (supplementary text, Fig. S2).
 369 We neglect other volatile species as they have generally lower concentrations and/or poorly con-
 370 strained solubility at Kīlauea (13, 43). We assume ideal gas behavior, and consider melt density

371 a function of pressure, temperature, and composition (29). Melt viscosity $\mu_l(z)$ is assumed
 372 to be a function of temperature and dissolved H₂O (31). The impact of bubbles on apparent
 373 magma viscosity depends upon the magnitude of capillary forces (30). For expected strain rates
 374 of $\sim 10^{-1} \text{ s}^{-1}$ associated with slow exchange flow in the conduit, bubbles less than ~ 10 cm
 375 across will increase apparent viscosity approximately according to $\mu = \mu_l / (1 - \bar{\phi})$ (Fig. S2),
 376 where $\bar{\phi}(z)$ is background magma porosity (30).

377 For conduit-reservoir resonance pressure at the base of the magma column is (20) $P'_0 =$
 378 $-\pi R_0^2 \langle u'_0 \rangle \sin(\theta_0) / C_r$, where C_r is the total storativity of the reservoir (reservoir volume change
 379 per unit pressure increase). The Halema'uma'u reservoir assumed here corresponds to a 'buoyancy-
 380 dominated' limit where reservoir pressure changes have a negligible effect on the magma
 381 column during VLPs (supplementary text) (9). Pressure at the top of the magma column is
 382 $P'_H = P_{ex} + \langle u'_H \rangle \sin(\theta_H) \bar{\rho}_H g$, where subscript H indicates evaluation at the top of the magma
 383 column and $P_{ex}(t)$ is external forcing (Fig. 1). This system is equivalent to a driven harmonic
 384 oscillator with frequency-dependent damping, and exhibits exponentially decaying oscillations
 385 in response to an impulsive forcing (Fig. S1). We find the resonant period and quality factor by
 386 solving numerically for the free response of the system (supplementary text).

387 VLP seismic event inversions

388 We assume a temporally fixed magma plumbing system geometry, except for lava lake radi-
 389 us and surface elevation which are interpolated from measurements (4, 32) (supplementary
 390 text). We choose reference fixed parameters based on previous constraints where available.
 391 Where minimal constraints are available, we test a range of values and select combinations that
 392 produce feasible inversion results over most of the timeline, as detailed in the supplementary
 393 text. We approximate the lava lake and conduit as cylinders, with a reference conduit radius of
 394 15 m and conduit dip of 90 degrees from horizontal (Fig. 1, Table S1).

395 We conduct inversions using the conduit-reservoir resonance model for the three free pa-
 396 rameters (temperature, X^{avg} , and ΔX) from the three target values for each VLP seismic event:
 397 conduit bottom (Halema'uma'u reservoir top) pressure, resonance period, and resonance quality
 398 factor (Fig. 2). We use an iterative nonlinear trust-region-reflective solver to find the combina-
 399 tion of free parameter values that minimizes misfit E

$$E = \frac{|\omega - \omega^*|}{\omega^*} + \frac{|Q - Q^*|}{Q^*} + \frac{|\bar{P}_0 - \bar{P}_0^*|}{\bar{P}_0^*} \quad (2)$$

400 where vertical bars indicate absolute value, asterisks indicates observed/target values, Q is res-
 401 onance quality factor, ω is resonance angular frequency, and \bar{P}_0 is magmastic pressure at the
 402 bottom of the conduit (top of the reservoir). To prevent unfeasible solutions, we impose bounds
 403 on the search space such that volatile mass fraction at all depths is between 0-7 wt% and tem-
 404 perature is between 900-1600 °C. In most cases there is an exact solution ($E = 0$), although for
 405 some VLP events (e.g., in 2009 and early 2010) exact solutions do not exist for the reference pa-
 406 rameters and the solver will find a local minimum instead. Grid searches indicate that the misfit

407 spaces are convex, so the solver is finding unique global minima and/or unique exact solutions.
408 Time-series analysis methods used to interpret inversions are detailed in the supplementary text.

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510 **Acknowledgments**

511 We thank the Hawaiian Volcano Observatory and Hawai‘i Volcanoes National Park. We thank
512 US Geological Survey staff including Kyle Anderson, Emily Montgomery-Brown, Matt Patrick,
513 Kendra Lynn, Ingrid Johanson, and Phillip Dawson for data and discussion of Kīlauea. We
514 thank Chao Liang and Eric Dunham for discussion of VLP events. We thank editor Blair
515 Schoene as well as Matt Haney and two anonymous reviewers for thorough and constructive

516 feedback. Any use of trade, firm, or product names is for descriptive purposes only and does
517 not imply endorsement by the U.S. Government.

518 **Funding**

519 National Science Foundation grant EAR-2036980 (LK)

520 **Author contributions**

521 Conceptualization: JC, LK

522 Methodology: JC

523 Writing: JC, LK

524 **Competing interests**

525 Authors declare that they have no competing interests

526 **Data and materials availability**

527 All data and codes are available at https://bitbucket.org/crozierjosh1/vlp_inversion_codes or
528 <https://doi.org/10.7910/DVN/1NNZTJ>. GNSS data can also be found at <http://geodesy.unr.edu/>,
529 the VLP seismic event catalog can also be found at <https://doi.org/10.7910/DVN/2UGFKE>, and
530 lava lake data can be found at <https://doi.org/10.5066/P9ULRPMM>.

531 **Supplementary Materials**

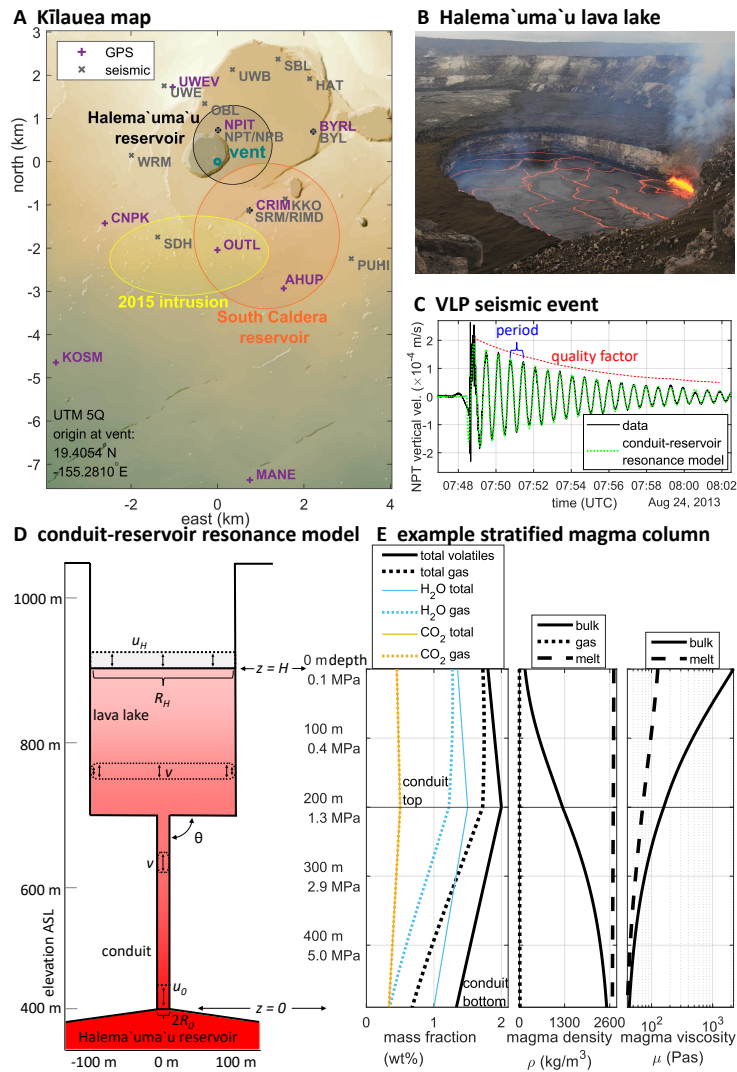
532 Materials and Methods

533 Supplementary Text

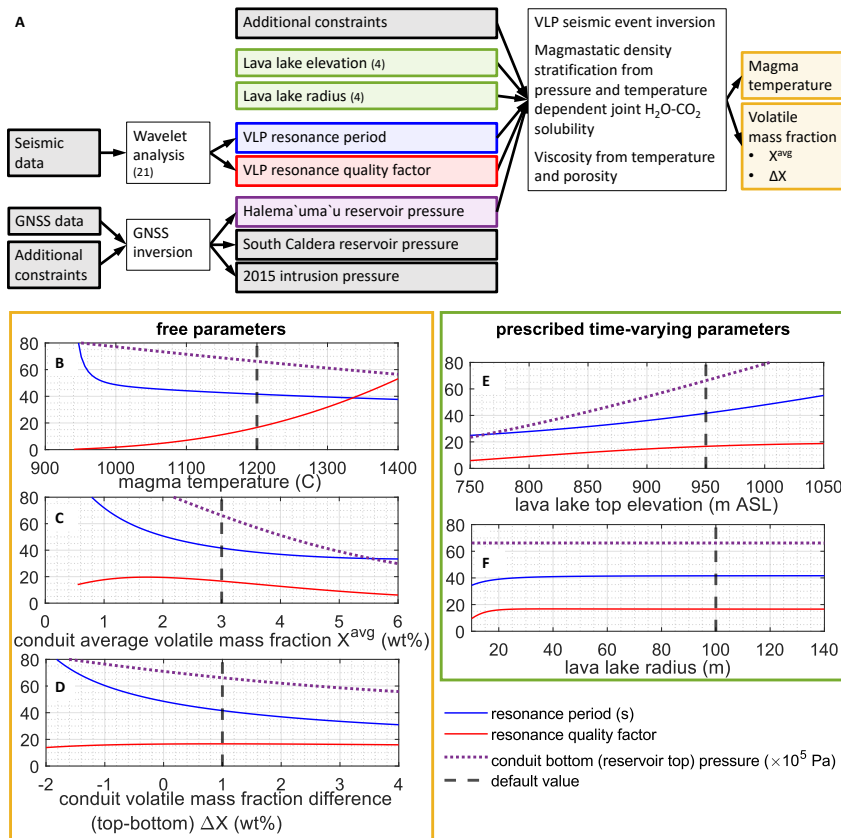
534 Figs. S1 to S10

535 Table S1

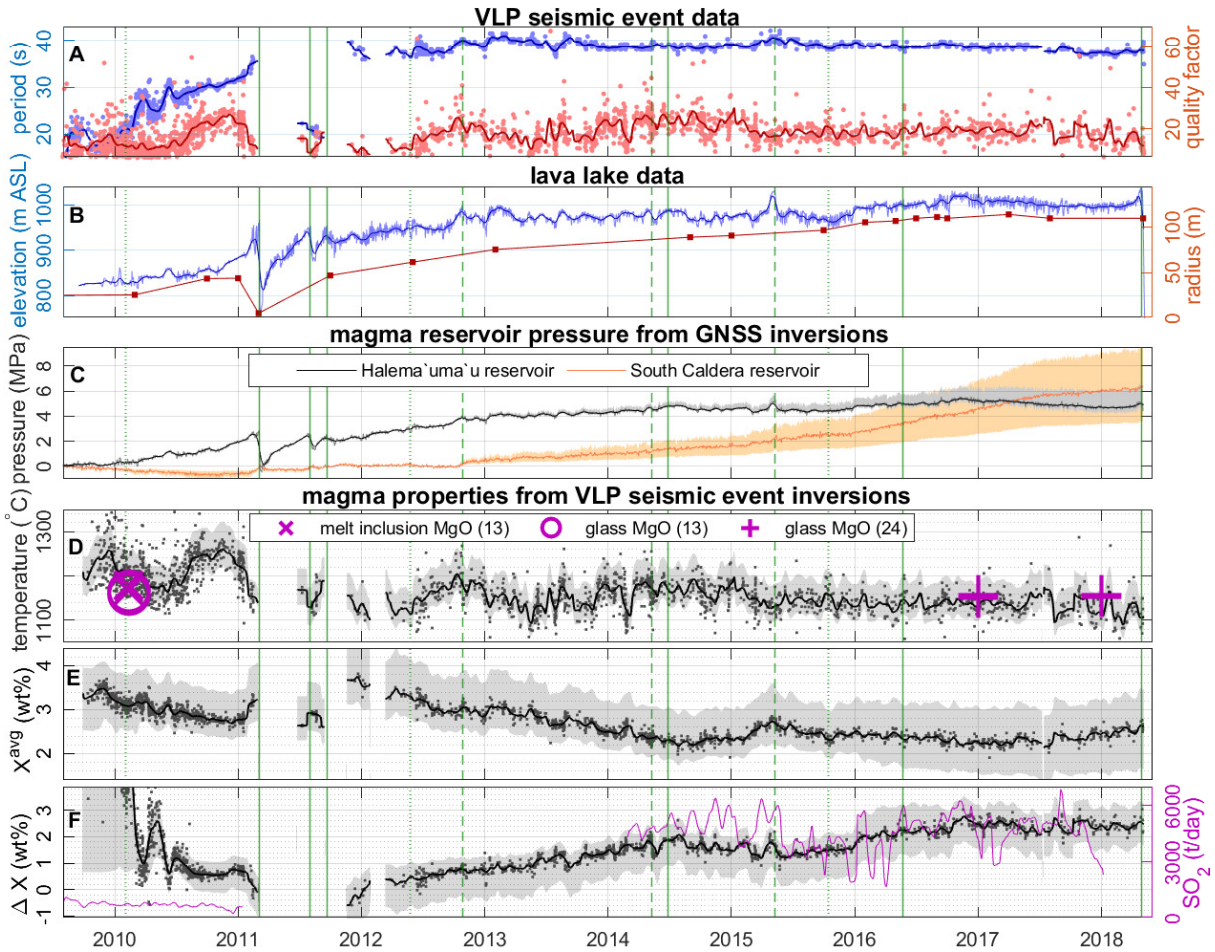
536 References (63-71)



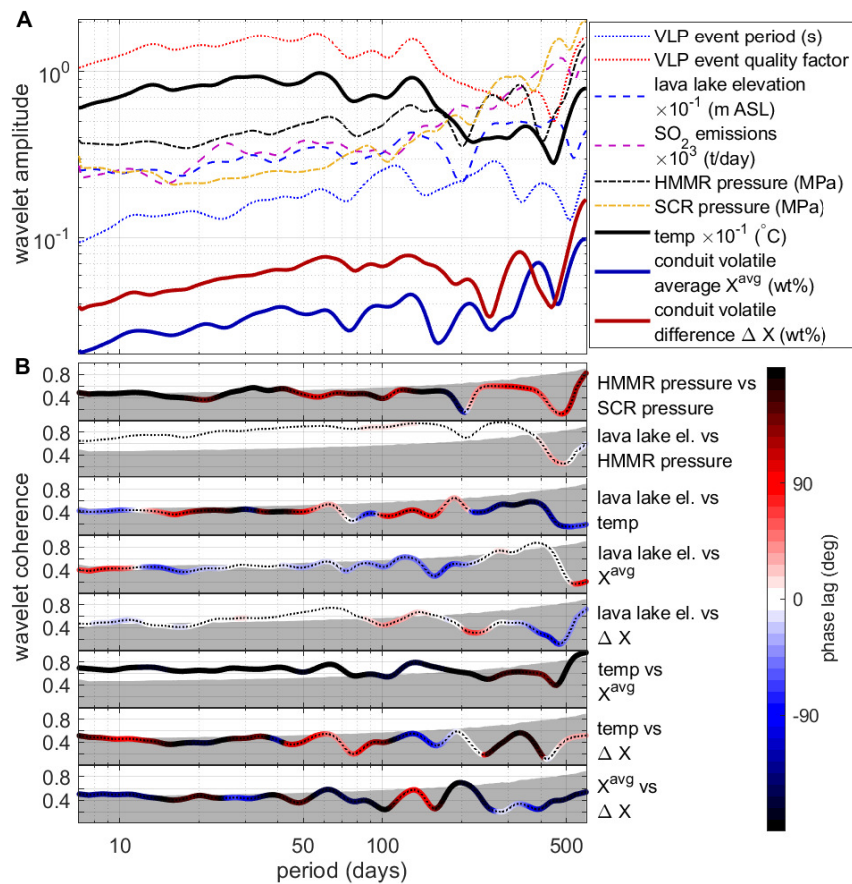
537 **Fig. 1. Kīlauea map and magma dynamics model.** (A) Map including the Halema'uma'u
 538 vent, inferred shallow magma storage zones, GNSS stations, and seismometers used in the
 539 VLP catalog (22). (B) Typical lava lake activity on Feb 13, 2017 (USGS). (C) Seismic wave-
 540 form from a VLP conduit-reservoir resonance event along with a model solution for reference
 541 fixed parameter inversion results forced with a Gaussian pressure perturbation (Fig. S1). (D)
 542 Conduit-reservoir resonance model with approximate 2018 magma system geometry; black ar-
 543 rows illustrate vertical sloshing of the stratified magma column. (E) Magmastic depth profiles
 544 from piecewise linear total (dissolved plus exsolved) volatile mass fractions at a uniform tem-
 545 perature of 1200 °C.



546 **Fig. 2. Inversion approach.** (A) Simplified flowchart of methods and data input/output. Additional
 547 constraints on GNSS inversions are from previous geodetic studies (11, 27, 57, 60). Additional
 548 constraints on VLP magma resonance inversions are from previous modeling (9), gravity
 549 data (37), and geochemical (gas and ejecta) data (13, 16, 24, 36). (B-F) Conduit-reservoir resonance
 550 period and quality factor, plus conduit bottom pressure, as a function of the parameters
 551 varied to fit Kīlauea VLP seismic and geodetic data. Variation in lava lake elevation and (assumed
 552 uniform) radius are prescribed from measurements (4, 32). Dashed black lines indicate
 553 default values used in the other plots.



554 **Fig. 3. Time series data and inversion results.** Inverted relative changes in magma prop-
 555 erties are from our reference fixed parameters (Fig. 1, Table S1). Dots represent individual
 556 VLP seismic events, bold lines are 30-day moving averages, while vertical green lines are East
 557 Rift Zone eruptions (solid), summit intrusions (dashed), and slow-slip events (dotted) (4). (A)
 558 VLP seismic event resonance period and quality factor (22). (B) Lava lake elevation and mean
 559 radius (4, 32) (C) GNSS inverted reservoir pressure changes, set to zero at the Mar 7, 2011 lava
 560 lake draining. Shaded areas indicate possible variation with different South Caldera reservoir
 561 geometries tested (supplementary text). (D) Inverted conduit magma temperature, with MgO
 562 thermometry for comparison (13, 24). The shaded area indicates possible variation with all
 563 fixed model parameter values tested (supplementary text). (E, F) Inverted conduit total volatile
 564 contents, with 30-day moving average SO₂ emissions for comparison (15, 16) and possible vari-
 565 ation shown in shaded areas. Values from 2009-early 2010 are unreliable due to exact solutions
 566 not being obtainable with the fixed parameters chosen.



567 **Fig. 4. Wavelet amplitude spectra and coherence.** (A) Amplitude spectra of resonance prop-
 568 erties (22), lava lake elevation (4, 32), SO₂ emissions (15, 16), GNSS inverted Halema'uma'u
 569 (HMMR) and South Caldera (SCR) reservoir pressures, and VLP magma resonance inverted
 570 magma properties. (B) Magnitude squared coherence colored by phase lag. The gray area is
 571 beneath the 95% significance threshold. Positive phase lags indicate that the second variable
 572 trails the first. Data before Dec 2011 were excluded from this analysis.

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Supplementary Material for

Evolving magma temperature and volatile contents over the 2008-2018 eruption of Kīlauea Volcano

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Supplementary Text

Conduit-reservoir model description

We consider a magmatic system consisting of a slowly convecting, vertically stratified column of fluid underlain by a reservoir in an elastic halfspace and overlain by a lava lake. Our model domain extends from the bottom of the conduit (or top of the reservoir) to the surface of the lava lake. To model VLP events, we separate the transient flow associated with small perturbations to this system from background dynamics. To first approximation, wave-like disturbances are rapid compared to background exchange flow so it suffices to consider a magmatic background state upon which small amplitude flow is superimposed (20). For sufficiently long period flow, we can neglect the compressibility of magma in the column and conduit wall elasticity. We adopt a coordinate system where z is the Cartesian direction parallel to the conduit/lava lake axis and r is the radial direction perpendicular to z , so the orientation of r and z is a function of conduit dip angle $\theta(z)$. Function arguments are omitted except where necessary. Linearized governing equations are derived for small amplitude uni-directional magma flow in this system using a perturbation approach,

$$[v(r, z, t), p(z, t), \rho(z, t)] = [0, \bar{p}(z), \bar{\rho}(z)] + [v'(r, z, t), p'(z, t), \rho'(z, t)], \quad (3)$$

where v is conduit-parallel fluid particle velocity (so the orientation of v is a function of conduit dip angle $\theta(z)$), p is pressure, ρ is magma density, overbar indicates background values, and prime indicates perturbations. We denote cross-sectional averaging as

$$\langle v \rangle(z, t) = \frac{2}{R^2} \int_0^R v(r, z, t) r dr \quad (4)$$

We also express motion in terms of cross-sectionally-averaged conduit-parallel fluid particle displacement $\langle u \rangle$,

$$\frac{d\langle u \rangle}{dt} = \langle v \rangle. \quad (5)$$

Magma density perturbation will result from advection of the background density profile

$$\rho = \langle u' \rangle \sin(\theta) \frac{d\bar{\rho}}{dz}. \quad (6)$$

Linearized conservation of momentum for perturbations is then given by

$$\frac{\partial v'}{\partial t} \bar{\rho} = \langle u' \rangle \sin(\theta) \frac{d\bar{\rho}}{dz} g - \sin(\theta) \frac{\partial p'}{\partial z} + \mu \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial v'}{\partial r} \right) \quad (7)$$

28 where μ is magma viscosity and g is gravitational acceleration. We assume a zero slip boundary
 29 condition along the magma column (conduit/lava lake) wall at radius $R(z)$

$$v'(z, R(z), t) = 0. \quad (8)$$

30 Viscous drag force can be determined from the shear stress at the magma column wall, so
 31 cross-sectionally averaging Eq. 7 gives

$$\frac{d\langle v' \rangle}{dt} \bar{\rho} = \langle u' \rangle \sin(\theta) \frac{d\bar{\rho}}{dz} g - \sin(\theta) \left[\frac{\partial p'}{\partial z} + \frac{2\mu}{R} \frac{\partial v'}{\partial r} \right]_R. \quad (9)$$

32 Incompressible linearized cross-sectionally averaged conservation of mass is

$$\langle u' \rangle = \langle u'_0 \rangle \frac{R_0^2}{R^2}, \quad (10)$$

33 where a zero subscript indicates evaluation at the bottom of the magma column (or top of the
 34 reservoir), e.g., $R_0 = R(z = 0)$, and subscript H indicates evaluation at the top of the magma
 35 column (or top of the lava lake).

36 We apply pressure perturbation boundary conditions at the top of the magma column and
 37 mass balance at the base. Neglecting fluid inertia and viscous dissipation in the reservoir due to
 38 long period forcing (20), linearized mass balance at the base of the magma column becomes a
 39 condition on basal pressure perturbation p'_0 ,

$$p'_0 = -\frac{\pi R_0^2 \langle u'_0 \rangle}{C_r} \sin(\theta_0), \quad (11)$$

40 where C_r is the total storativity (injected magma volume per unit pressure increase) of the
 41 reservoir,

$$C_r = (\beta_m + \beta_r) V_r, \quad (12)$$

42 where V_r is reservoir volume, $\beta_m = \frac{1}{\rho} \frac{d\bar{\rho}_r}{dp_r}$ is effective magma compressibility in the reservoir,
 43 and $\beta_r = \frac{1}{V_r} \frac{dV_r}{dp_r}$ is the elastic reservoir compressibility. For a spherical reservoir $\beta_r = \frac{3}{4G}$, where
 44 G is the host rock elastic shear modulus (63). Linearized pressure perturbation at the top of the
 45 magma column p'_H is a function of external forcing pressure $P_{ex}(t)$ and the displaced magma
 46 mass at the free surface,

$$p'_H = P_{ex} + \langle u'_H \rangle \sin(\theta_H) \bar{\rho}_H g = p'_{ex} + \langle u'_0 \rangle \sin(\theta_H) \frac{R_0^2}{R_H^2} \bar{\rho}_H g. \quad (13)$$

47 Integrating momentum (Eq. 9) in the z -direction over magma column height H and substi-
 48 tuting in conservation of mass (Eq. 10) and the boundary conditions (Eq. 11 and 13) gives

$$\begin{aligned} & \frac{d^2 \langle u'_0 \rangle}{dt^2} R_0^2 \int_0^H \frac{1}{\sin(\theta)} \frac{\bar{\rho}}{R^2} dz = \\ & \langle u'_0 \rangle R_0^2 \left(g \left(\int_0^H \frac{d\bar{\rho}}{dz} \frac{1}{R^2} dz - \bar{\rho}_H \frac{1}{R_H^2} \sin(\theta_H) \right) - \frac{\pi}{C_r} \sin \theta_0 \right) + \int_0^H \frac{1}{\sin(\theta)} \frac{2\mu}{R} \left[\frac{\partial v'}{\partial r} \right]_R dz - P_{ex}. \end{aligned} \quad (14)$$

49 Conduit-reservoir model solution

50 We assume a periodic pressure gradient with angular frequency ω and amplitude f , to focus
51 on the fundamental eigenmode of the system (the conduit-reservoir oscillation)

$$\frac{\partial p'}{\partial z} = f e^{i\omega t}. \quad (15)$$

52 Velocity can then be expressed analytically (64)

$$v' = \frac{f e^{i\omega t}}{i\omega \bar{\rho}} \left(1 - \frac{J_0(r\alpha)}{J_0(R\alpha)} \right), \quad (16)$$

53 where J_n is a Bessel function of the first kind and order n , and α is

$$\alpha = \sqrt{\frac{\omega \bar{\rho}}{\mu}} i^{3/2} \quad (17)$$

54 with $i = \sqrt{-1}$. Shear strain rate at the conduit/lava lake wall is then

$$\left. \frac{\partial v'}{\partial r} \right|_R = \frac{f e^{i\omega t}}{i\omega \bar{\rho}} \left(\frac{\alpha J_1(R\alpha)}{J_0(R\alpha)} \right), \quad (18)$$

55 and cross-sectionally averaged velocity $\langle v' \rangle$ is

$$\langle v' \rangle = \frac{f e^{i\omega t}}{i\omega \bar{\rho}} \left(1 - \frac{2}{R\alpha} \frac{J_1(R\alpha)}{J_0(R\alpha)} \right). \quad (19)$$

56 Substituting Eq. 19 into Eq. 18 and simplifying with the Bessel function recurrence relation

57 $J_{n+1}(x) = \frac{2n}{x} J_n(x) - J_{n-1}(x)$ yields

$$\left. \frac{\partial v'}{\partial r} \right|_R = -\langle v' \rangle \frac{\alpha J_1(R\alpha)}{J_2(R\alpha)}. \quad (20)$$

58 Substituting Eq. 20, 5, and 10 into Eq. 14 and taking the real part finally gives

$$\begin{aligned} & \frac{d^2 \langle u'_0 \rangle}{dt^2} R_0^2 \int_0^H \frac{1}{\sin(\theta)} \frac{\bar{\rho}}{R^2} dz = -P_{ex} \\ & + \langle u'_0 \rangle R_0^2 \left(g \left(\int_0^H \frac{d\bar{\rho}}{dz} \frac{1}{R^2} dz - \bar{\rho}_H \frac{1}{R_H^2} \sin(\theta_H) \right) - C_r^{-1} \pi \sin(\theta_0) \right) \\ & - \frac{d \langle u'_0 \rangle}{dt} 2R_0^2 \mathbf{Re} \left[\int_0^H \frac{1}{\sin(\theta)} \frac{\mu}{R^3} \frac{\alpha J_1(R\alpha)}{J_2(R\alpha)} dz \right]. \end{aligned} \quad (21)$$

59 This equation can be solved in the frequency domain for a given time function of P_{ex} . Examples
60 of such solutions are given in Fig. 1 and Fig. S1.

61 To study the natural response of the conduit-reservoir oscillation we set the top external
 62 forcing pressure P_{ex} in Eq. 21 to zero (rendering forcing an initial condition), which gives a
 63 homogeneous damped harmonic oscillator equation

$$c_1 \frac{d^2 \langle u'_0 \rangle}{dt^2} + c_2 \frac{d \langle u'_0 \rangle}{dt} + c_3 \langle u'_0 \rangle = 0. \quad (22)$$

64 In equation 22, c_1 scales the magnitude of the inertial term for the oscillator

$$c_1 = R_0^2 \int_0^H \frac{1}{\sin(\theta)} \frac{\bar{\rho}}{R^2} dz. \quad (23)$$

65 c_2 scales the viscous damping term

$$c_2 = 2R_0^2 \mathbf{Re} \left[\int_0^H \frac{1}{\sin(\theta)} \frac{\mu}{R^3} \frac{\alpha J_1(R\alpha)}{J_2(R\alpha)} dz \right], \quad (24)$$

66 and is a function of ω through α . c_3 scales the restoring force term (gravity and reservoir
 67 storativity),

$$c_3 = -R_0^2 \left(g \left(\int_0^H \frac{d\bar{\rho}}{dz} \frac{1}{R^2} dz - \frac{\bar{\rho}_H}{R_H^2} \sin(\theta_H) \right) - C_r^{-1} \pi \sin(\theta_0) \right). \quad (25)$$

68 Equation 22 has a general solution of the form

$$\langle u'_0 \rangle(t) = \langle u'_0 \rangle(t=0) e^{(\lambda+i\omega)t}, \quad (26)$$

69 with initial amplitude $\langle u'_0 \rangle(t=0)$ set by the external pressure perturbation, temporal exponen-
 70 tial decay rate constant

$$\lambda = \frac{c_2}{2c_1}, \quad (27)$$

71 and natural angular frequency

$$\omega = \sqrt{\frac{c_3}{c_1} - \left(\frac{c_2}{2c_1} \right)^2} = \sqrt{\omega_u^2 - \lambda^2}, \quad (28)$$

72 where undamped (inviscid) natural angular frequency $\omega_u = \sqrt{c_3/c_1}$. Since c_2 is a function of
 73 ω , Eq. 28 must be solved implicitly for ω , which then may be used to calculate λ from Eq. 27.
 74 Quality factor Q gives the ratio of energy stored to energy lost per oscillation cycle,

$$Q = \frac{\omega}{2\lambda}. \quad (29)$$

75 Conduit-reservoir model analytical solutions under simplified conditions

76 To gain more insight into these equations, we examine a simplified scenario that permits
 77 a concise analytical solution. We consider a vertical cylindrical magma column with uniform
 78 magma viscosity. We assume a linear magma density gradient between $\bar{\rho}_0$ and $\bar{\rho}_H$, alternately
 79 characterized by the vertically averaged density $\bar{\rho}_{avg} = (\bar{\rho}_H + \bar{\rho}_0)/2$ and the vertical density
 80 difference $\Delta\bar{\rho} = \bar{\rho}_H - \bar{\rho}_0$. We assume fully developed (Poiseuille) flow, which will provide an
 81 upper bound on viscous damping. This simplified scenario is similar to those considered in (21)
 82 and in the reduced conduit-reservoir eigenmode model of (20). In this scenario, the inertial
 83 scale factor reduces to

$$c_1 = H\bar{\rho}_{avg}, \quad (30)$$

84 the viscous damping scale factor reduces to Poiseuille drag

$$c_2 = H\frac{8\mu}{R^2}, \quad (31)$$

85 and the restoring force scale factor reduces to

$$c_3 = \bar{\rho}_0g + \pi R^2 C_r^{-1}. \quad (32)$$

86 This yields an exponential decay rate of

$$\lambda = \frac{4\mu}{R^2\bar{\rho}_{avg}}, \quad (33)$$

87 a natural angular frequency of

$$\omega = \sqrt{\frac{g(\bar{\rho}_H - \Delta\bar{\rho}) + \pi R^2 C_r^{-1}}{H\bar{\rho}_{avg}} - \frac{16\mu^2}{R^4\bar{\rho}_{avg}^2}}, \quad (34)$$

88 and a quality factor of

$$Q = \frac{R^2\bar{\rho}_{avg}}{8\mu} \sqrt{\frac{g(\bar{\rho}_H - \Delta\bar{\rho}) + \pi R^2 C_r^{-1}}{H\bar{\rho}_{avg}} - \frac{16\mu^2}{R^4\bar{\rho}_{avg}^2}}. \quad (35)$$

89 The natural frequency of flow that is not fully developed, as will be the case during Kīlauea VLP
 90 events (20), will be between the natural frequency of fully developed flow and the undamped
 91 natural frequency (for which Q is not defined)

$$\omega_u = \sqrt{\frac{g(\bar{\rho}_H - \Delta\bar{\rho}) + \pi R^2 C_r^{-1}}{H\bar{\rho}_{avg}}}. \quad (36)$$

92 This simplified scenario permits an easy examination of the relative importance of restoring
 93 forces from gravity and reservoir storativity for the Kīlauea magma system geometry. The

94 compressibility of the ellipsoidal Halema‘uma‘u reservoir geometry (11) is $\approx 2.5 \times 10^{-10} \text{ Pa}^{-1}$.
 95 Magma compressibility in the reservoir could range from 10^{-9} to 10^{-10} Pa^{-1} (9), from which
 96 Eq. 12 gives reservoir storativity of $\sim 1\text{-}5 \text{ m}^3/\text{Pa}$. For a conduit radius of 10-20 m (9), the
 97 reservoir storativity restoring force term in Eq. 32 will range from $\sim 60\text{-}300 \text{ N/m}^3$. The density
 98 difference across the conduit will likely be at least $\sim 1000 \text{ kg/m}^3$ (9). The gravity restoring force
 99 term in Eq. 32 will thus be at least $\sim 10^4 \text{ N/m}^3$, which is an order of magnitude larger than the
 100 reservoir storativity term. This is consistent with a similar analysis in (20).

101 Stratified magma properties

102 We prescribe piecewise linear depth profiles of magma temperature and total (dissolved plus
 103 exsolved) volatile contents, parameterized by their value at the bottom of the conduit, top of the
 104 conduit, and top of the lava lake. Density and viscosity are then calculated from these magma
 105 properties. We consider both CO_2 and H_2O , but do not explicitly treat other volatiles as their
 106 solubility and/or abundance is poorly constrained.

107 We approximate the background pressure profiles as magmastic

$$\bar{p}(z) = \bar{P}_{atm} + \int_z^H \bar{\rho}(y)gdy, \quad (37)$$

108 where atmospheric pressure $\bar{P}_{atm} = 10^5 \text{ Pa}$. Exchange flow could result in sub-magmastic
 109 pressures (28), but this is not well constrained by data used here. Background bulk magma
 110 density is given by

$$\bar{\rho}(z) = \left(\frac{\bar{n}_g(z)}{\bar{\rho}_g(z)} + \frac{1 - \bar{n}_g(z)}{\bar{\rho}_l(z)} \right)^{-1}. \quad (38)$$

111 Where \bar{n}_g is background gas mass fraction, $\bar{\rho}_l(z)$ is background gas density, and $\bar{\rho}_l(z)$ is back-
 112 ground melt density. We calculate background melt density as a function of pressure, tem-
 113 perature, and composition using the model of (29) with average Halema‘uma‘u melt inclusion
 114 compositions from Table 7 in (13). We use the ideal gas law for background gas density:

$$\bar{\rho}_g(z) = \frac{\bar{p}(z)(\bar{n}_{H_2O}(z)M_{H_2O} + \bar{n}_{CO_2}(z)M_{CO_2})}{R_g T} \quad (39)$$

115 where \bar{n}_m and M_m are the background exsolved gas mass fraction and molar mass of volatile
 116 species m , T is temperature, and R_g is the ideal gas constant.

117 To obtain exsolved gas mass fractions from total (dissolved plus exsolved) volatile mass
 118 fractions (\bar{X}_{H_2O} , \bar{X}_{CO_2}), we interpolate pre-computed volatile solubility from the model of (25).
 119 These give equilibrium H_2O and CO_2 solubility as a function of pressure and H_2O gas molar
 120 fraction, again using average Halema‘uma‘u melt inclusion compositions from (13). The ac-
 121 curacy of the chemical equilibrium assumption depends on the rate of magma ascent/descent
 122 relative to the rate of volatile diffusion in/out of bubbles. Estimated lava lake upwelling ve-
 123 locities of 0.15-0.3 m/s would yield magma ascent timescales in the lava lake on the order of

124 hours (53). H₂O and CO₂ diffusivity are highly dependent on temperature and H₂O contents,
 125 but should be on the order of 10⁻⁹ to 10⁻¹¹ m²/s in the shallow Kīlauea magma system (65).
 126 This could correspond to chemical diffusion timescales from minutes to hours for typical bub-
 127 ble spacing of 10⁻⁵ to 10⁻³ m, and potentially longer in a regime dominated by isolated large
 128 bubble slugs (66).

129 We calculate melt viscosity $\mu_l(z)$ as a function of temperature and dissolved H₂O from the
 130 model of (31), again using the average Kīlauea glass composition from Table 7 in (13). Crystal
 131 contents (67) will increase bulk magma viscosity, but we neglect this given the relatively low
 132 crystal contents of Halema‘uma‘u magma (13).

133 The effect bubbles have on bulk magma viscosity depends upon the flow regime (30). For
 134 oscillatory flows, this is governed by the dynamic capillary number, which is the ratio between
 135 the timescale over which bubbles relax to spherical shapes and the timescale over which changes
 136 in shear deformation occur: $C_d = \frac{\mu_l R_b \ddot{\epsilon}}{\Gamma \dot{\epsilon}}$. For $C_d < 1$ bubbles will act as obstacles to flow and
 137 increase bulk magma viscosity, whereas for $C_d > 1$ bubbles will act as weak regions that deform
 138 preferentially and reduce bulk magma viscosity. Bubble radii R_b in effusive Hawaiian eruptions
 139 are on the order of 10⁻⁴ to 10⁻³ m, although there will likely be some lateral variability (28)
 140 and bubble slugs with widths up to the conduit width could occasionally be present (33, 66).
 141 However, very large bubble slugs would break-up and/or ascend on the order of minutes (33, 68),
 142 and since Strombolian-type bubble bursts only occur intermittently (33) we assume that the
 143 conduit and lava lake are free of such large bubble slugs most of the time. We additionally note
 144 that if a bursting bubble slug triggers VLP resonance, that slug would not be present during the
 145 following resonance. Melt viscosity μ_l will be on the order of 10¹ to 10² Pa·s (31). Surface
 146 tension Γ will be on the order of 10⁻¹ N/m (69). The mean strain rate ratio $\ddot{\epsilon}/\dot{\epsilon}$ for a sinusoidal
 147 velocity will be approximately $2\pi/T$, so on the order of 10⁻¹ s⁻¹ for these VLP events. C_d will
 148 then generally be on the order of 10⁻³ to 10⁻¹. We thus use the $C_d < 1$ capillary number model
 149 from (30) for background bulk magma viscosity,

$$\mu(z) = \frac{\mu_l(z)}{1 - \bar{\phi}(z)}, \quad (40)$$

150 where $\bar{\phi}(z)$ is background magma porosity, $\bar{\phi} = (\bar{\rho}_l - \bar{\rho})/(\bar{\rho}_l - \bar{\rho}_g)$. This relation becomes
 151 inaccurate as porosity approaches 1, such as in foam layers that might build up near the lava
 152 lake surface. However, we will show in the next section that the lava lake contributes negligibly
 153 to viscous damping during conduit-reservoir resonance.

154 Fig. S2 shows the effects of temperature and total (dissolved plus exsolved) volatile contents
 155 on magma properties.

156 Conduit-reservoir model exploration

157 We consider model parameters that are plausible for the Kīlauea magma plumbing system.
 158 We approximate the lava lake geometry as a vertical cylinder in all of our simulations. This is
 159 justified for the case of Kīlauea since at both times when the lava lake fully drained its geometry

160 was roughly cylindrical (4), and we also found that using a conical frustum approximation to the
161 lava lake geometry produced values of period and quality factor that differed from a cylindrical
162 geometry by less than 1%. There are no direct constraints on conduit geometry except for
163 limited observations from the times when the lava lake drained fully, where it appears that the
164 top of the conduit is appreciably smaller than the base of the lava lake (4). Previous inversion
165 by (9) of isolated Halema‘uma‘u conduit-reservoir VLP events with a model similar to ours
166 assuming a cylindrical conduit indicates a steeply dipping conduit with a most likely radius of
167 10-20 m. We consider conduit geometries consisting of either cylinders or conical frustums,
168 and allow the conduit to dip at an angle θ from vertical.

169 While some previous VLP seismic inversions have inferred a source geometry of intersect-
170 ing dikes (21, 33), an ellipsoidal reservoir is consistent with the collapse geometry observed in
171 2018 (11), with other geodetic inversions (7, 10, 58, 59), and previous work combining model-
172 ing with VLP seismic inversions (9). We thus adopt the ellipsoidal reservoir geometry and rock
173 shear modulus found by (11) as our reference scenario. Simulations for our assumed Kīlauea
174 magma system geometry verify that reservoir storativity has a negligible impact on resonant
175 period and quality factor in this system, consistent with the analysis above (conduit-reservoir
176 model analytical solutions under simplified conditions) and in (20). We thus fix the compress-
177 ibility of magma in the reservoir to $5 \times 10^{-10} \text{ Pa}^{-1}$.

178 Fig. S9 shows the effects of various magma system geometries and magma properties on
179 resonant period, quality factor, and conduit bottom magmastatic pressure load (equal to pressure
180 at the top of Halema‘uma‘u reservoir). For comparison, Fig. S10 shows simulations where
181 magma density and viscosity are directly prescribed following piecewise linear depth functions.
182 In this case lava lake elevation and magma properties in the lava lake do not appreciably effect
183 resonant period or quality factor (Fig. S10) because the much larger cross-sectional area of
184 the lava lake relative to the conduit means that the viscous damping, inertial, and gravitational
185 terms are comparatively minimal in the lava lake. However, in the volcanologically informed
186 background state lava lake elevation and magma properties in the lava lake do affect period and
187 quality factor. These parameters change the magmastatic pressure load on the conduit, thus
188 changing volatile solubility and gas density. This illustrates one important advantage of using
189 the volcanologically informed background state model.

190 GNSS inversions for reservoir pressure change

191 Table S1 details our assumed reservoir geometry. We use daily GNSS solutions due to
192 significant noise in higher frequency GNSS, and the instrumental drift in tilt-meter data that can
193 be significant at timescales of months or longer. We correct GNSS displacements for the steady
194 background flank slip motion using the multi-component (dikes and décollement) dislocation
195 source model of (57) which consists of slip along low-angle normal faults as well as opening
196 and strike-slip motion along segments of the east rift zone (Fig. S3).

197 We find that Fourier domain first-order topography corrections (70) change inverted pres-
198 sures by less than 1%, so we do not include them for consistency with the south flank motion

199 corrections which were derived without topography. For each time t we use a linear least-
 200 squares inversion to jointly solve for pressure changes in the two/three reservoirs that best fit
 201 the observed displacements $U_{j,k}$ for the east, north and vertical ($k = E, N, Z$) components of
 202 $j = 1 : m$ available stations,

$$U_{j,k}(t) = G_{j,k}^{HMM} \Delta P^{HMM}(t) + G_{j,k}^{SCR} \Delta P^{SCR}(t) + G_{j,k}^{INT} \Delta P^{INT}(t), \quad (41)$$

203 where G^{HMM} , G^{SCR} , and G^{INT} and are halfspace quasistatic elastic Green's functions for the
 204 Halema'uma'u reservoir, South Caldera reservoir, and 2015 intrusion respectively (62), and
 205 ΔP^{HMM} , ΔP^{SCR} , and ΔP^{INT} are pressure changes.

206 VLP event inversions for magma properties

207 Table S1 lists reference values of all fixed parameters used for these inversions. We linearly
 208 interpolate between lava lake surface elevation and surface area measurements in (4, 32) to di-
 209 rectly prescribe lava lake surface elevation and effective lava lake radius (assuming a circular
 210 lava lake surface) at the time of each VLP event. We do not interpolate lava lake bottom eleva-
 211 tion since there are only two measurements in 2011 and 2018 (4). To obtain the target conduit
 212 bottom pressure at the time of each VLP event, we add an assumed baseline pressure to our
 213 geodetically inverted Halema'uma'u reservoir pressure changes and linearly interpolate to the
 214 time of VLP events (see section below on inversions with different fixed parameters).

215 We note that exact solutions to the data do not imply zero uncertainty, as there is uncertainty
 216 in the data. Uncertainty in VLP event ω and Q depends upon factors such as the signal/noise
 217 ratio of each event, and is highly variable (22). We use only the more robustly resolved events,
 218 for which uncertainty in ω is ~ 2 -4% of the inverted values of ω . Uncertainty in Q is more diffi-
 219 cult to robustly quantify, but we estimate it to be ~ 5 -50%. Uncertainty in GNSS displacements
 220 is ~ 0.001 m (61); $\sim 0.1\%$ of the total displacements from 2008-2018 (~ 1 m) and $\sim 10\%$ of the
 221 maximum daily displacements (~ 0.01 m). Uncertainty in inverted reservoir pressure changes
 222 (as a percentage) will be of a similar order of magnitude to the uncertainty in GNSS data. Un-
 223 certainty in lava lake elevation measurements is 1-5 m (4, 32). Additional uncertainty is also
 224 present in reservoir pressure and lava lake elevation from interpolating these data to the time of
 225 each VLP event.

226 Uncertainty in Q by far dominates the total data uncertainty. Since temperature in our in-
 227 versions results is primarily a function of Q , uncertainty in temperature will be dominated by
 228 uncertainty in Q and may be up to ~ 100 °C. Uncertainty in Q also ends up being the largest
 229 contributor to uncertainty in inverted total volatile contents, since variation in inverted tempera-
 230 ture induced by uncertainty in Q effects gas density and induces uncertainty of up to ~ 0.5 wt%
 231 in volatile contents. We thus expect that noise in Q contributes to much of the scatter in all
 232 inverted magma properties.

233 Time-series analysis

234 To mitigate noise in the time-series of inversion results induced by data error, we calculate
 235 moving averages with a 30-day triangular weighted moving window. This window was chosen

236 to smooth much of the apparent scatter while preserving trends over timescales of weeks or
237 longer.

238 To produce uniformly sampled data for frequency analysis, we first linearly interpolate
239 all data sets at 1 hr increments. We use continuous wavelet transforms with Morlet wavelets
240 (since our time series are non-stationary) for each individual data set, and continuous wavelet
241 magnitude-squared coherence and cross spectra between each pair of data sets (e.g., Fig. S7).
242 We then calculate mean values across the timeline at each frequency in a continuous wavelet
243 transform or continuous wavelet coherence to estimate the overall spectrum or coherence. To
244 obtain the overall cross spectrum we use a weighted mean based on the magnitude-squared co-
245 herence at each time and frequency, which ensures that the overall values more strongly reflect
246 the times where signals are more coherent.

247 To estimate 95% significance thresholds for coherence, we generate 10000 pairs of synthetic
248 Gaussian white noise and compute coherence between each pair following the methods above
249 (71). The 95% threshold for each frequency is then taken to be the 95th quantile of coherence at
250 that frequency (i.e., there is only a 5% chance that values above this threshold could be random
251 noise rather than coherent signals).

252 Inversions with different fixed parameters

253 Here we only consider variation in the fixed parameters that are most poorly constrained
254 and/or that have the largest effect on inverted magma properties, and we focus on the effects
255 of changing each parameter in isolation. Fig. S5 shows these effects relative to the reference
256 values in Table S1.

257 We find empirically that South Caldera reservoir centroid depth and aspect ratio (height/width)
258 have nearly identical impacts on inverted Halema‘uma‘u reservoir pressure changes (hence on
259 inverted magma properties), so we only show the former. Decreasing either parameter causes
260 the inversions to assign more of the long-term deformation to the Halema‘uma‘u reservoir. We
261 find that the South Caldera reservoir needs to be relatively deep and/or vertically elongated to
262 produce time-series of pressure in the Halema‘uma‘u reservoir that are consistent with observed
263 lava lake elevation. For example, we show approximate bounds on such magmastatic pressure
264 changes in Fig. S4 that were calculated assuming lower and upper bounds on average magma
265 column densities of 1000 and 2700 kg/m³. Since previous studies have found either vertically
266 shortened or spherical South Caldera reservoir geometries, we assume a reference spherical ge-
267 ometry. The volume of a spherical reservoir does not significantly effect ground deformation
268 patterns (63) and thus does not impact inverted Halema‘uma‘u reservoir pressure, so we fix the
269 South Caldera reservoir volume to 20 km³ (58). Previous studies find centroid elevations rang-
270 ing from -2 to -4 km ASL (7, 10, 58, 59), so we choose a reference of -3 km ASL. Decreasing
271 this to -4 km ASL has a negligible impact on temperature, increases X^{avg} by ~ 0.3 wt% by the
272 end of the timeline, and decreases ΔX by ~ 0.4 wt% by the end of the timeline. Increasing this
273 to -2 km ASL has a negligible impact on temperature, decreases X^{avg} by ~ 1 wt% by the end
274 of the timeline, and increases ΔX by ~ 0.6 wt% by the end of the timeline. We also note that

275 this shallower South Caldera reservoir causes a strong trend in Halema‘uma‘u reservoir pres-
276 sure such that these inversions are not able to exactly fit both pressure and ω after 2016 without
277 invoking magma densities in excess of the pure melt density at the base of the conduit.

278 Minimal direct constraints exist on possible values for H₂O/CO₂ mass ratio in the shallow
279 magma system since many of the volatiles (particularly CO₂) are exsolved. Estimates of the
280 volatile mass ratio in primitive/parent magma at depth vary but are typically around 1 (36). Sig-
281 nificant outgassing of CO₂ at depth results in estimated Halema‘uma‘u gas emission H₂O/CO₂
282 mass ratios that are highly variable but up to 30, and Halema‘uma‘u melt inclusion and glass
283 compositions show a wide range of H₂O/CO₂ mass ratios (13, 36). We thus chose an interme-
284 diate reference H₂O/CO₂ mass ratio of 3 (or 1 wt% H₂O-to- 4.7×10^3 ppm CO₂). Decreasing
285 the mass ratio to 1 uniformly increases temperature by ~ 10 °C, uniformly increases X^{avg} by
286 ~ 0.8 wt%, and uniformly increases ΔX by ~ 0.6 wt%. Increasing the mass ratio to 20 uni-
287 formly decreases temperature by ~ 10 °C, uniformly decreases X^{avg} by ~ 0.4 wt%, and uni-
288 formly decreases ΔX by ~ 0.4 wt%. We set the baseline Halema‘uma‘u reservoir top pressure
289 relative to the time of the Mar 7, 2011 lava lake draining (Fig. S4). Bounds on this baseline can
290 be obtained by considering magmatic pressure from feasible conduit average magma densi-
291 ties (say 400-2600 kg/m³). However, we find that many baseline pressures that are feasible at
292 this particular time would require unfeasibly high or low magma densities in some part of the
293 conduit at other times. We choose a reference baseline Halema‘uma‘u reservoir top pressure
294 of 2.3 MPa that corresponds to an average magma column density of 800 kg/m³ at the time
295 of the Mar 7, 2011 lava lake draining; this produces feasible densities/volatile mass fractions
296 over all of the timeline after 2010. Decreasing baseline pressure to 2.0 MPa (average magma
297 column density of 700 kg/m³) uniformly increases temperature by ~ 10 °C, uniformly increases
298 X^{avg} by ~ 0.2 wt%, and uniformly decreases ΔX by ~ 0.2 wt%. Increasing baseline pressure
299 to 2.6 MPa (average magma column density of 900 kg/m³) uniformly decreases temperature by
300 ~ 10 °C, uniformly decreases X^{avg} by ~ 0.2 wt%, and uniformly increases ΔX by ~ 0.2 wt%.

301 We choose a reference conduit radius of 15 m, which produces temperatures generally con-
302 sistent with or less than geochemically inferred Halema‘uma‘u reservoir values (which we as-
303 sume represent an approximate upper bound on plausible conduit temperatures) (7, 13, 24).
304 Decreasing the conduit radius to 5 m uniformly increases temperature by ~ 50 °C, uniformly
305 decreases X^{avg} by ~ 0.1 wt%, and has a negligible impact on ΔX . Increasing the conduit radius
306 to 25 m uniformly decreases temperature by ~ 30 °C, uniformly increases X^{avg} by ~ 0.1 wt%,
307 and has a negligible impact on ΔX .

308 We choose a reference conduit length of 290 m, which is consistent with a vertical connec-
309 tion between the Halema‘uma‘u reservoir geometry we adopt from (11) and our assumed lava
310 lake base elevation of 700 m ASL. We explore the effect of varying conduit length by varying
311 lava lake base elevation, noting that varying the reservoir top elevation would have a roughly
312 similar effect (Fig. S9). Decreasing the conduit length to 190 m uniformly decreases tem-
313 perature by ~ 20 °C, uniformly increases X^{avg} by ~ 0.2 wt%, and uniformly increases ΔX by
314 ~ 0.5 wt%. Increasing the conduit length to 390 m uniformly increases temperature

315 by ~ 20 °C, has a negligible impact on X^{avg} (except in the earliest part of the timeline), and
316 uniformly decreases ΔX by ~ 1 wt%.

317 Available continuous gravity data suggest that the top of the lava lake is a foam with poros-
318 ity of 92-96%, varying on timescales of hours with episodic ‘gas-pistoning’ events (4, 37).
319 We chose a reference lava lake top volatile contents of 1.8 wt%, corresponding to a porosity
320 of $\sim 93\%$. Decreasing lava lake top volatile contents to 1 wt% uniformly increases temper-
321 ature by ~ 10 °C, uniformly increases X^{avg} by ~ 0.8 wt%, and uniformly increases ΔX by
322 ~ 0.4 wt%. Increasing lava lake top volatile contents to 2.6 wt% uniformly decreases temper-
323 ature by ~ 10 °C, uniformly decreases X^{avg} by ~ 0.4 wt%, and uniformly decreases ΔX by
324 ~ 0.4 wt%.

325 Additional coherence and phase lag calculations

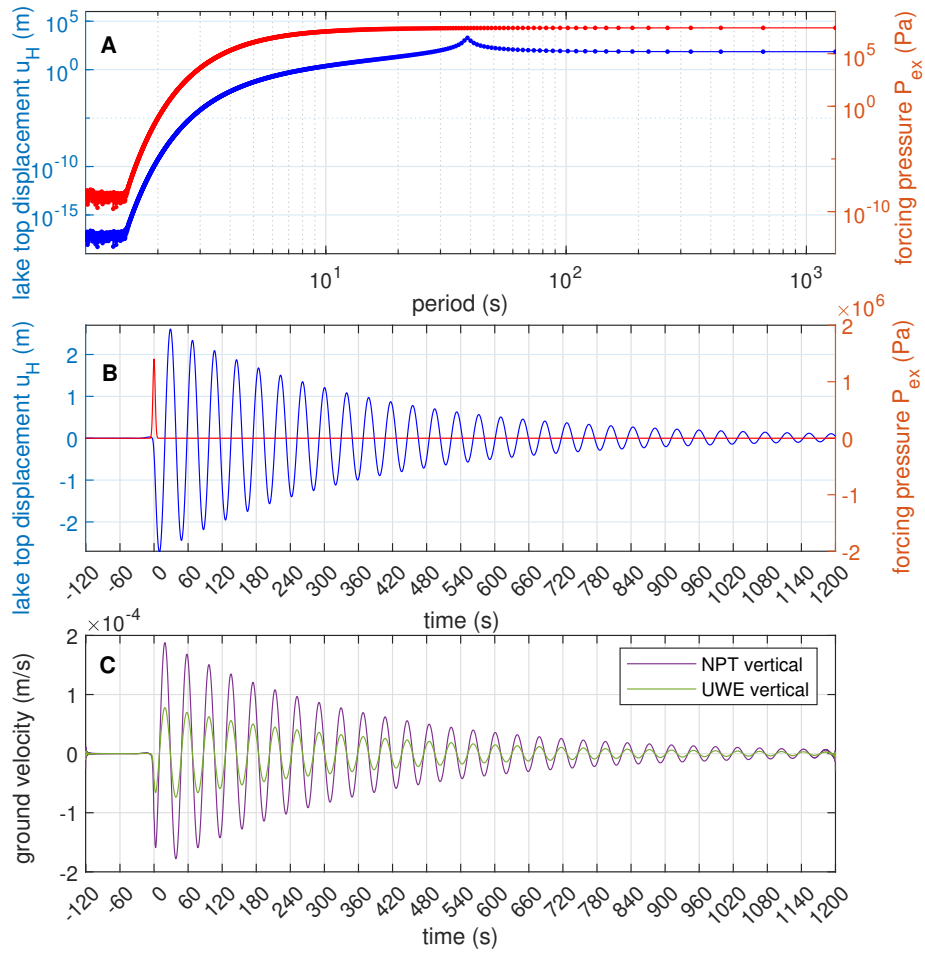
326 Fig. S6 and Fig. S7 show additional coherence and phase lag plots.

327 Inversions for direct values of magma density and viscosity

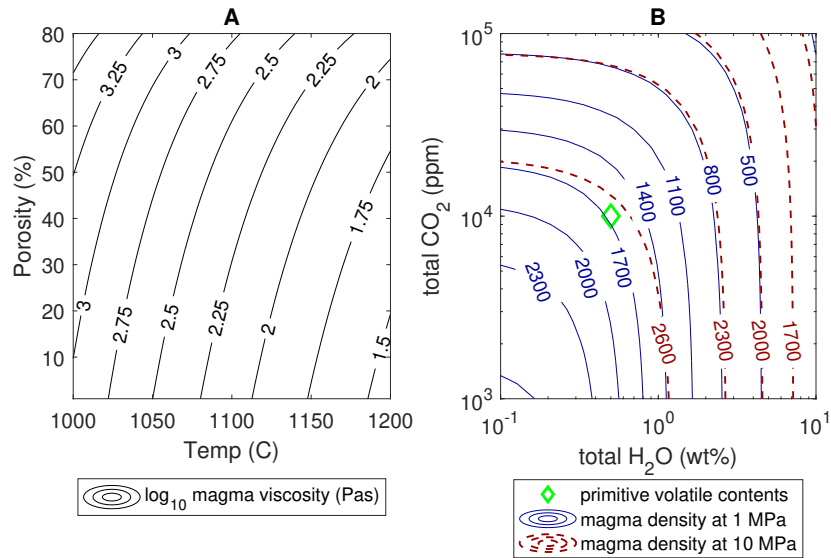
328 We show results from inversions for magma density and viscosity in Fig. S8. This provides
329 context for the inversions for volatile contents and temperature we focus on in the main text,
330 and facilitates comparison with inversions for isolated VLP events in (9). For these inversions
331 we assume a uniform magma viscosity and a fixed magma density at the top of the lava lake,
332 analogous to the assumptions we made in temperature and volatile content inversions. The three
333 free parameters are then: (1) magma density at the conduit top, (2) magma density at the conduit
334 bottom, and (3) uniform magma viscosity. Density is also shown in Fig. S9 as the average value
335 in the conduit and the difference between the top and bottom of the conduit.

336 There is over an order of magnitude of variation in inverted viscosity on timescales ranging
337 from days to years. For the majority of the 2009-2018 timespan, magma viscosity exhibits a
338 clear inverse relationship with Q (Fig. S8). This is consistent with the strong impact of viscosity
339 on Q (Fig. S10). Part of the large scatter in viscosity is likely related to noise in the estimates
340 of Q (22).

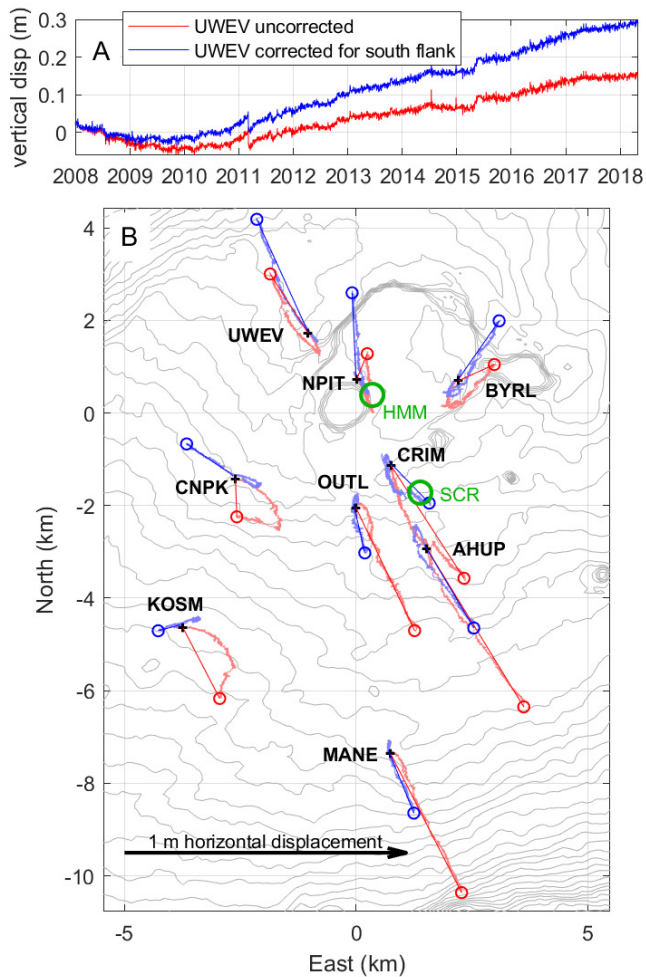
341 Conduit averaged magma density and conduit magma density difference both roughly track
342 lava lake elevation and inverted reservoir pressure. A positive relation between conduit average
343 magma density and lava lake elevation/reservoir pressure is expected since changing lava lake
344 elevation shifts the magma column up or down. Vertical translation of the magma column could
345 also explain the variations observed in density difference if the density gradient is more gradual
346 at greater depths (i.e., nonlinear), which is expected unless volatile contents increase signifi-
347 cantly with depth. This dependence of density upon upward/downward shifting of the magma
348 column highlights another important advantage of using volcanologically informed background
349 states to infer changes in properties of interest such as volatile contents.



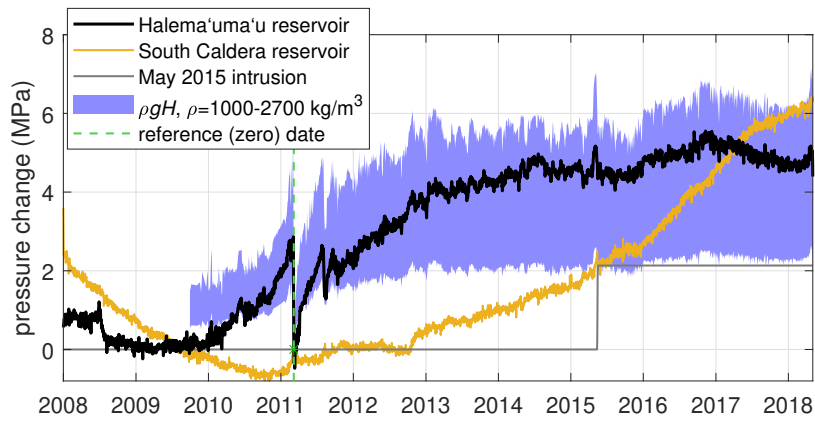
350 **Fig. S1.** Example solution to the conduit-reservoir magma resonance model for a Gaussian
 351 pressure perturbation with amplitude of 1.4 MPa and variance of 4 s applied to the magma
 352 column shown in main text Fig. 1. (A) Lava lake top displacement and forcing pressure in the
 353 frequency domain. (B) Lava lake top displacement and forcing pressure in the time domain.
 354 (C) Vertical ground velocities at the locations of two nearby seismometers.



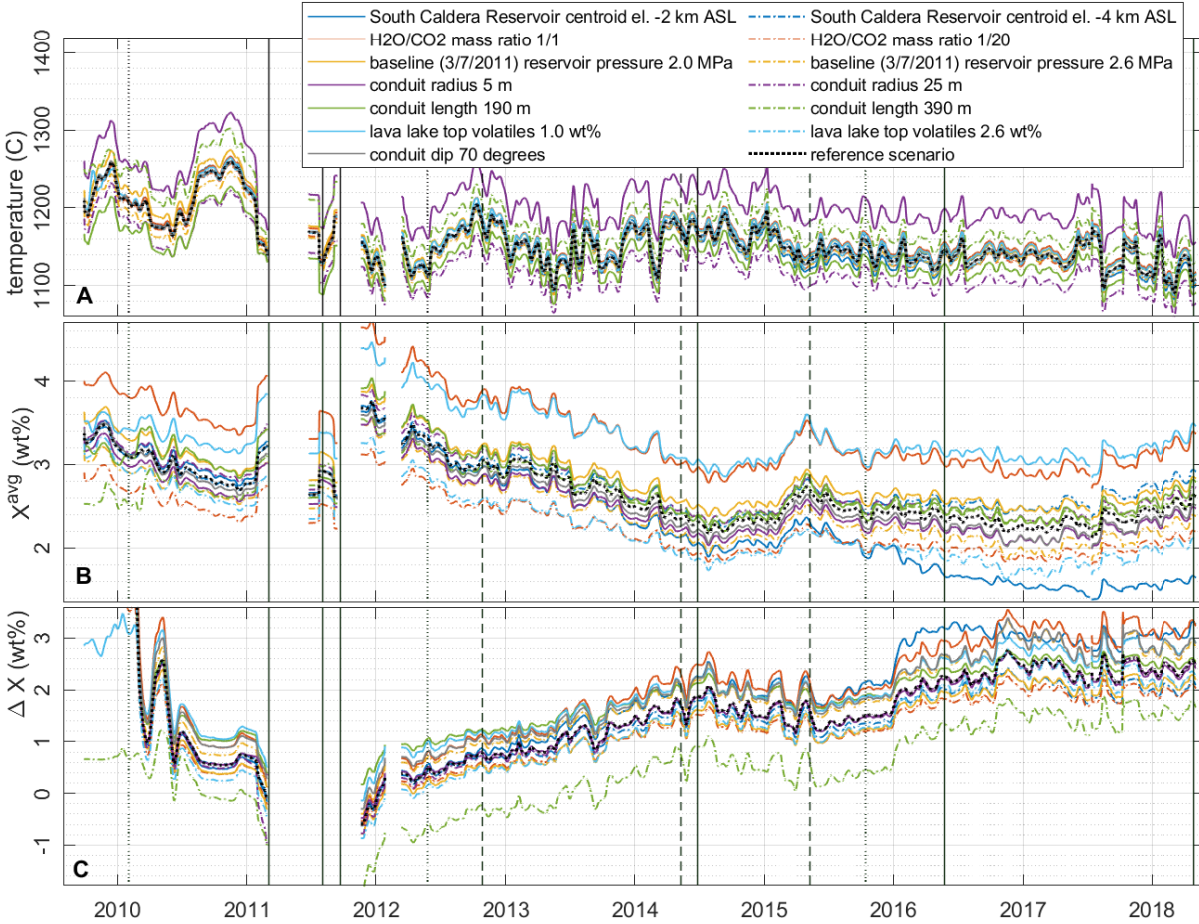
355 **Fig. S2.** (A) Variation of magma viscosity with temperature and porosity, adopted from (22).
 356 (B) Variation of magma density with H_2O and CO_2 contents at two pressures (1 and 10 MPa
 357 correspond to magmastatic depths of 40-100 m and 0.4-1 km respectively) and a temperature
 358 of 1100 °C. The density of pure melt is $\sim 2650 \text{ kg/m}^3$. Estimated primitive (or ‘parent’) magma
 359 total volatile contents from (36).



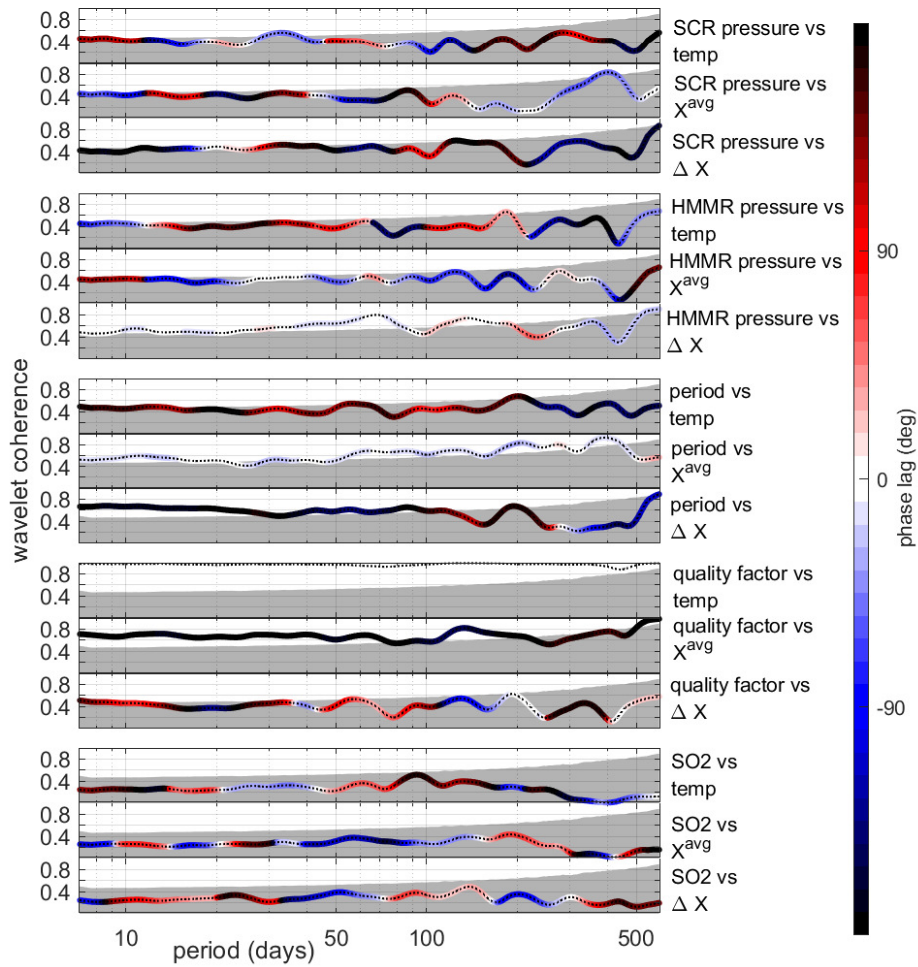
360 **Fig. S3.** (A) Vertical ground displacement at GNSS station UWEV. (B) Map of GNSS stations
 361 and horizontal ground displacements from 2008-2018. Red lines are corrected for flank motion
 362 and blue are uncorrected. Lighter red and blue lines show the GNSS displacement over time
 363 from Jan 1, 2008 (black plus symbols) to May 1, 2018 (red and blue circles). Straight red and
 364 blue lines show the net 2008 to 2018 displacement vectors. The inferred centroid locations of
 365 the Halema'uma'u (HMM) and South Caldera (SCR) reservoirs are shown by green circles.
 366 UTM zone 5Q.



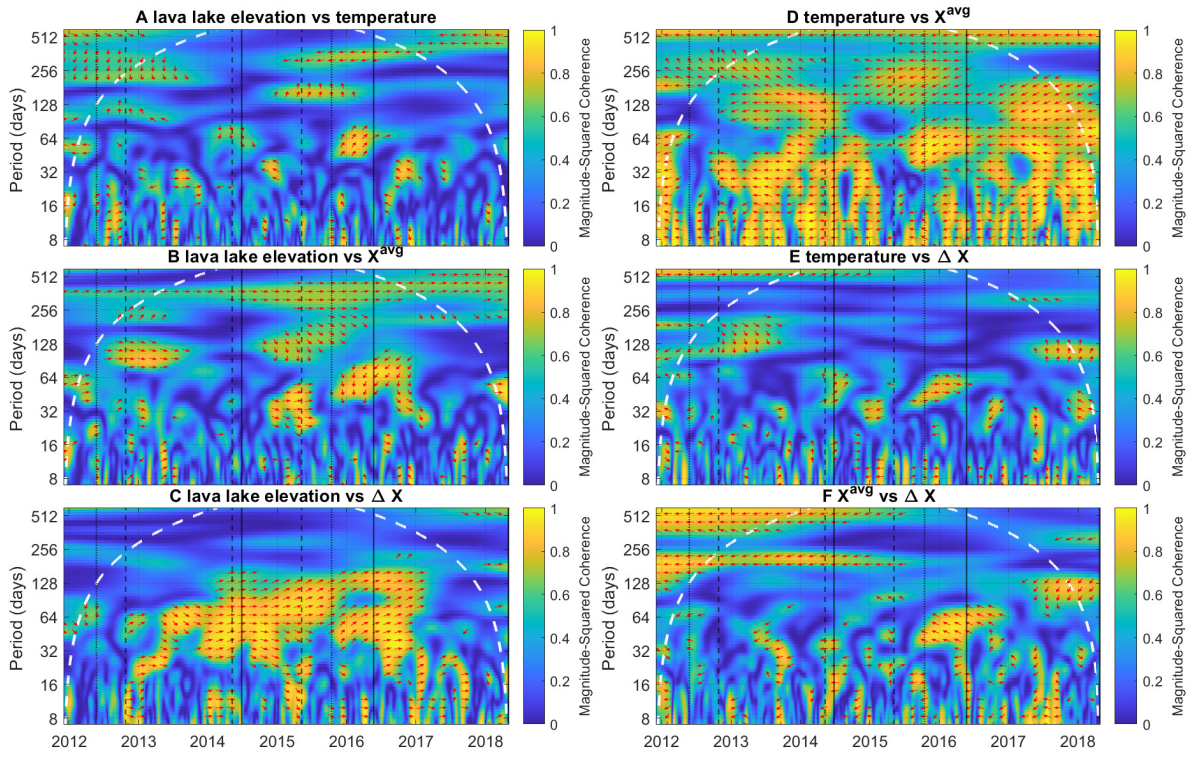
367 **Fig. S4.** Joint GNSS inversions for pressure change over 2008-2018 compared to magmatic
 368 pressure changes calculated from lava lake elevation data (4, 32) with assumed average magma
 369 column densities ρ of 1000 and 2700 kg/m³ (approximate lower and upper bounds).



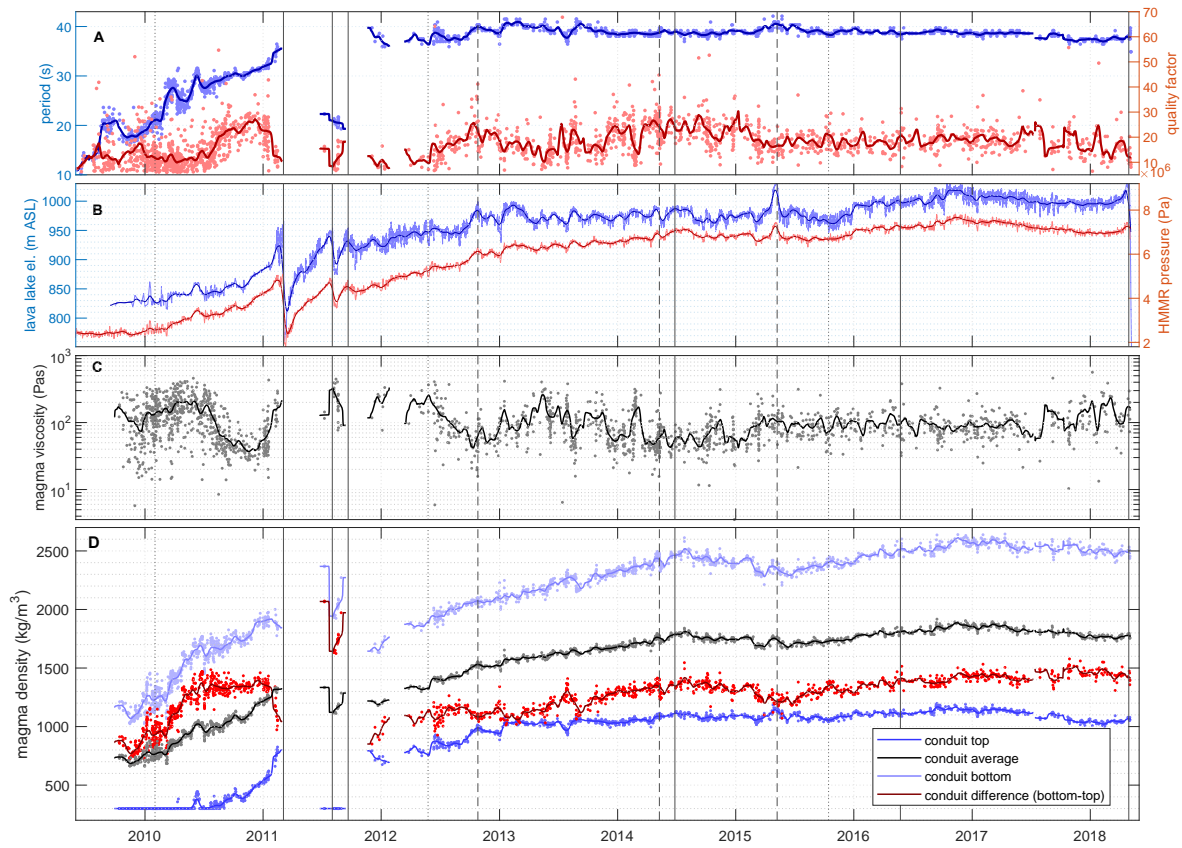
370 **Fig. S5.** Inversions with different fixed parameter values. The legend indicates the one fixed
 371 parameter value changed to produce each solid colored line; all other fixed parameters are held
 372 equal to the reference values from Table S1. Vertical black lines are East Rift Zone eruptions
 373 (solid), summit intrusions (dashed), and slow-slip events (dotted) (4). Values from 2009-early
 374 2010 are unreliable due to exact solutions not being obtainable with the fixed parameter combi-
 375 nations shown.



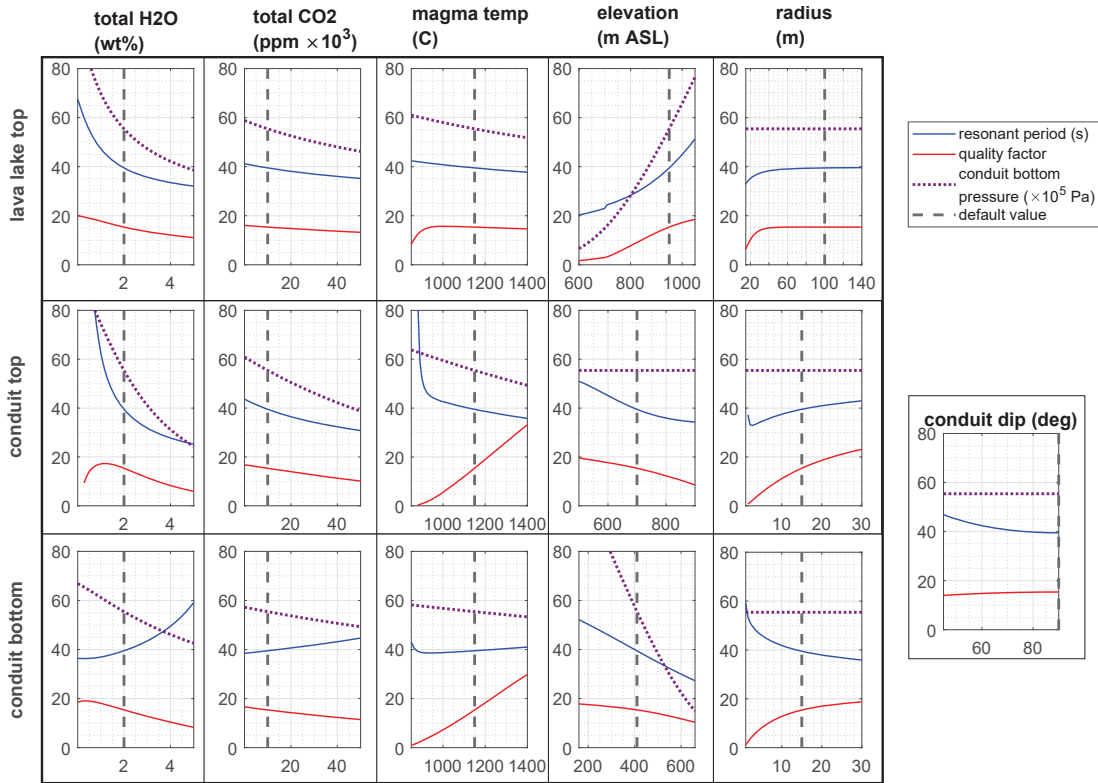
376 **Fig. S6.** Magnitude squared coherence colored by phase lag. The gray area is beneath the 95%
 377 significance threshold. Positive phase lags indicate that the second variable trails the first. Data
 378 before Dec 2011 were excluded from this analysis.



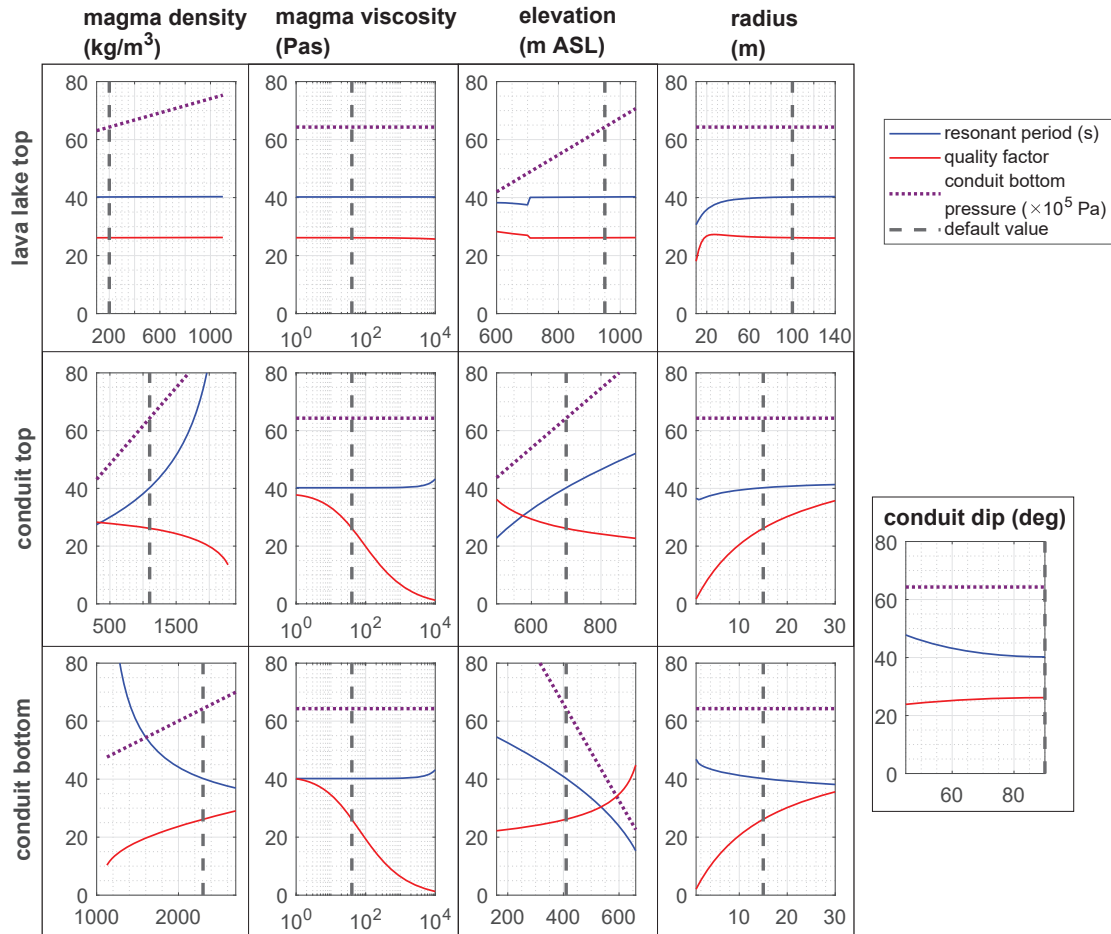
379 **Fig. S7.** Spectral coherence and phase lags between lava lake elevation (4, 32)
 380 based magma properties. Red arrows indicate the direction of phase lag where coherence is
 381 greater than 0.5; right indicates in-phase (positive correlation), left indicates 180 degrees out
 382 of phase (negative correlation), and up or down indicates 90 degrees out of phase. The white
 383 region in 2011 was excluded due to limited data. Dashed white lines indicate the region of edge
 384 influence. Vertical black lines are East Rift Zone eruptions (solid), summit intrusions (dashed),
 385 and slow-slip events (dotted) (4).



386 **Fig. S8.** Inverted relative changes in magma properties for our reference fixed parameter val-
 387 ues (Table S1) without parameterization in terms of temperature and volatile contents. Dots
 388 represent individual VLP events, bold lines are 30-day moving averages, while vertical black
 389 lines are East Rift Zone eruptions (solid), summit intrusions (dashed), and slow-slip events (dot-
 390 ted) (4). **(A)** VLP period and quality factor (22). **(B)** Lava lake elevation (4,32) and geodetically
 391 inverted reservoir pressure changes, relative to the time of the Mar 7, 2011, lava lake draining.
 392 **(C, D)** Inverted magma properties.



393 **Fig. S9.** Predicted variation in resonance period, quality factor, and pressure at the bottom
 394 of the conduit (or top of the reservoir) due to varying model parameters in isolation. Dashed
 395 black lines indicate the default value of each parameter used to make the other plots. We do not
 396 show parameters related to reservoir storativity (reservoir shape, host rock shear modulus, and
 397 magma compressibility in the reservoir) since they have a negligible impact on these simula-
 398 tions. We note that conduit bottom elevation is the same as reservoir top elevation and that we
 399 have assumed a cylindrical lava lake.



400 **Fig. S10.** Predicted variation in resonance period, quality factor, and pressure at the bottom
 401 of the conduit (or top of the reservoir) due to varying model parameters in isolation without
 402 parameterization in terms of temperature and volatile contents. Dashed black lines indicate the
 403 default value of each parameter used to make the other plots. We do not show parameters related
 404 to reservoir storativity (reservoir shape, host rock shear modulus, and magma compressibility
 405 in the reservoir) since they have a negligible impact on these simulations. We note that conduit
 406 bottom elevation is the same as reservoir top elevation and that we have assumed a cylindrical
 407 lava lake.

	parameter	default value	units
Conduit + lava lake geometry			
	lava lake top elevation	prescribed from (4, 32)	m ASL
	lava lake bottom elevation	700	m ASL
	conduit bottom elevation	410	m ASL
H	total (conduit + lava lake) length	calculated	m
R_H	lava lake radius	prescribed from (4, 32)	m
	conduit top radius	15	m
R_0	conduit bottom radius	15	m
θ	conduit dip (from horizontal)	90	degrees
Magma reservoirs			
G	rock shear modulus	3.08	GPa
	Halema'uma'u reservoir geometry	fixed from (11)	
	2015 intrusion geometry	fixed from (27)	
	South Caldera reservoir centroid elevation	-3000	m ASL
	South Caldera reservoir centroid latitude	19.3900	degrees
	South Caldera reservoir centroid longitude	-155.2710	degrees
	South Caldera reservoir volume	20	km ³
	South Caldera reservoir aspect ratio	1	
Magma properties			
β_m	Halema'uma'u reservoir magma compressibility	5×10^{-10}	Pa ⁻¹
	H ₂ O/CO ₂ mass ratio	3	
	lava lake top volatiles	1.8	wt%
	melt composition	fixed from (13)	
ρ_l	melt density	calculated from (29)	
μ_l	melt viscosity	calculated from (31)	
	H ₂ O-CO ₂ solubility	calculated from (25)	
Other			
P_{atm}	atmospheric pressure	10 ⁵	Pa
p_0	baseline (Mar 7, 2011) conduit bottom pressure	2.3	MPa
g	gravitational acceleration	9.81	m/s ²
R_g	ideal gas constant	8.314	Jkg ⁻¹ mol ⁻¹

408 **Table S1.** Reference model parameter values